FYS 3610

EXERCISES WEEK 47

EXERCISE 1

- a) Discuss the usage of the Reynolds number as an indicator of whether the frozen-in-flux concept is valid or not. Why does the frozen-in-flux concept (ideal MHD) break down locally near a reconnection site.
- b) Point out what are the major differences between the Sweet-Parker and Petschek model of magnetic reconnection. What allows Petschek reconnection to proceed at such a high rate compared to a Sweet-Parker diffusion region?

EXERCISE 2

- a) Describe the main concepts of the Cowley-Lockwood model of driving polar cap convection by transient reconnection?
- b) Assume that IMF B_X and B_Y are both zero. Assume that IMF Bz has been steadily positive for several hours, that the polar cap is contracted, and that polar cap convection has calmed. I.e. the solar wind-magnetosphere coupling has been poor for a long time. Then IMF suddenly turns negative and maintains negative. Describe the ionospheric effect of this sudden transition in IMF. Put emphasis on flow generation, polar cap boundary motion, and how a new equilibrium will be achieved. About how long will it take until equilibrium of balanced magetopause and magnetotail reconnection has accomplished?
- c) Starting out again with a quiet situation. The polar cap is circular with a radius of r = 1500 km. The ionospheric magnetic field is $B = 5 \times 10^{-5}$ T (roughly perpendicular to the Earth surface). Estimate the total amount of open flux.
- d) Assume that a reconnection pulse lasting 5 minutes adds 10% more open flux to the dayside. Estimate the voltage along the dayside reconnection line associated with this pulse.
- e) Magnetopause reconnection between closed magnetic field lines and the draped IMF, takes place with a steady voltage of 100 kV for an extended period, while reconnection in the cross-tail current sheet takes place with a voltage of 20 kV. Describe the changes in the magnetosphere-ionosphere system that will take place

while this situation persists. Initially the polar cap is a circle of radius 1500 km what radius will it have after 30 min? What do we call such an interval? (The ionospheric magnetic field is 5×10^{-5} T).

EXERCISE 3

Figure 1 illustrates crossing through the auroral cusp by the FAST satellite.

- a) Describe the typical energy of incoming electrons and ions (pitch angle zero is down along the magnetic field line).
- b) What are the most likely source for these particles.
- c) Note the energy dispersion of ions which is taken as a characteristic of cusp precipitation. Can you provide an explanation for the energy dispersed ions (decreasing with latitude)?
- d) dBz in panel b) of the FAST plot is magnetic field variation perpendicular to the orbital plane. Introduce position and direction of Birkeland current sheets from the positive and negative slopes. Where are the Birkeland currents located with respect to the auroral form. Comment your answer.

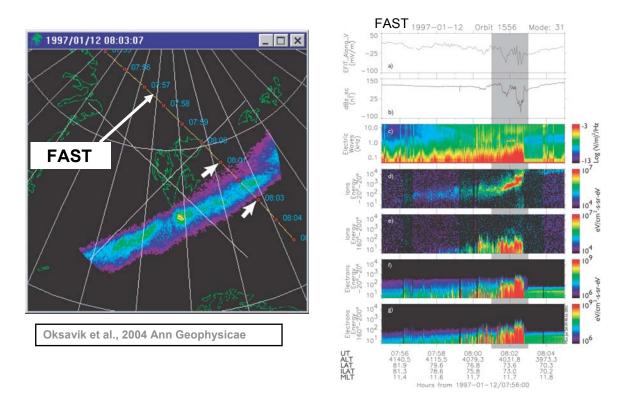


Figure 1 630.0 nm All-sky image from Ny-Ålesund. The straight yellow line indicates the FAST passage over through it, from which data are presented on the right.

Exercise 14.8 Kivelson & Russell