

# Climate of planet Earth 1



# **Climate of planet Earth 1**

- **Basic principles**
- **Main climate drivers**
- **Climate in the geological past**

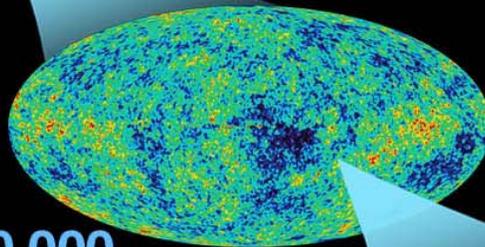
# **Facts about planet Earth**

**DAWN  
OF  
TIME**



**tiny fraction  
of a second**

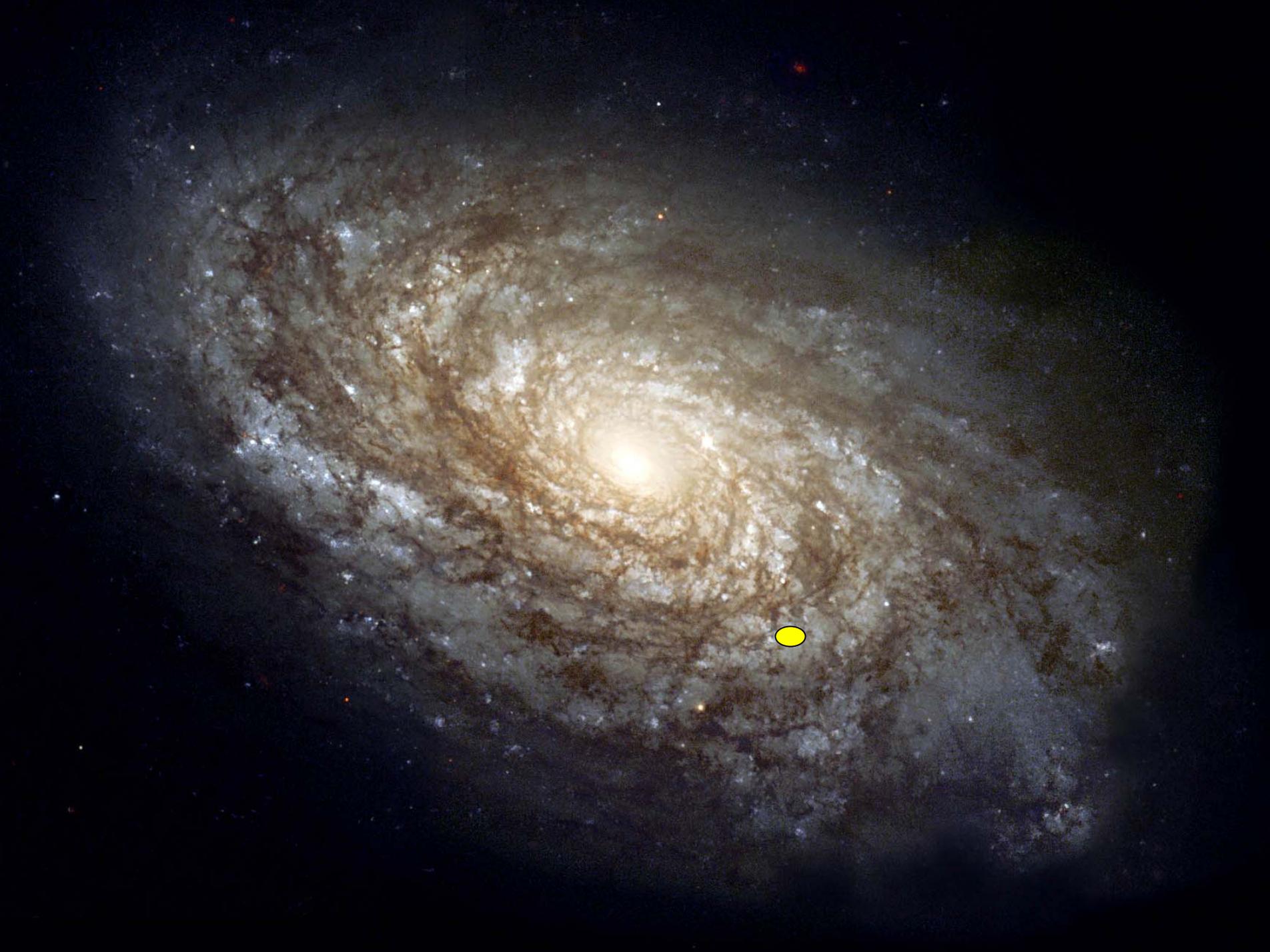
**inflation**

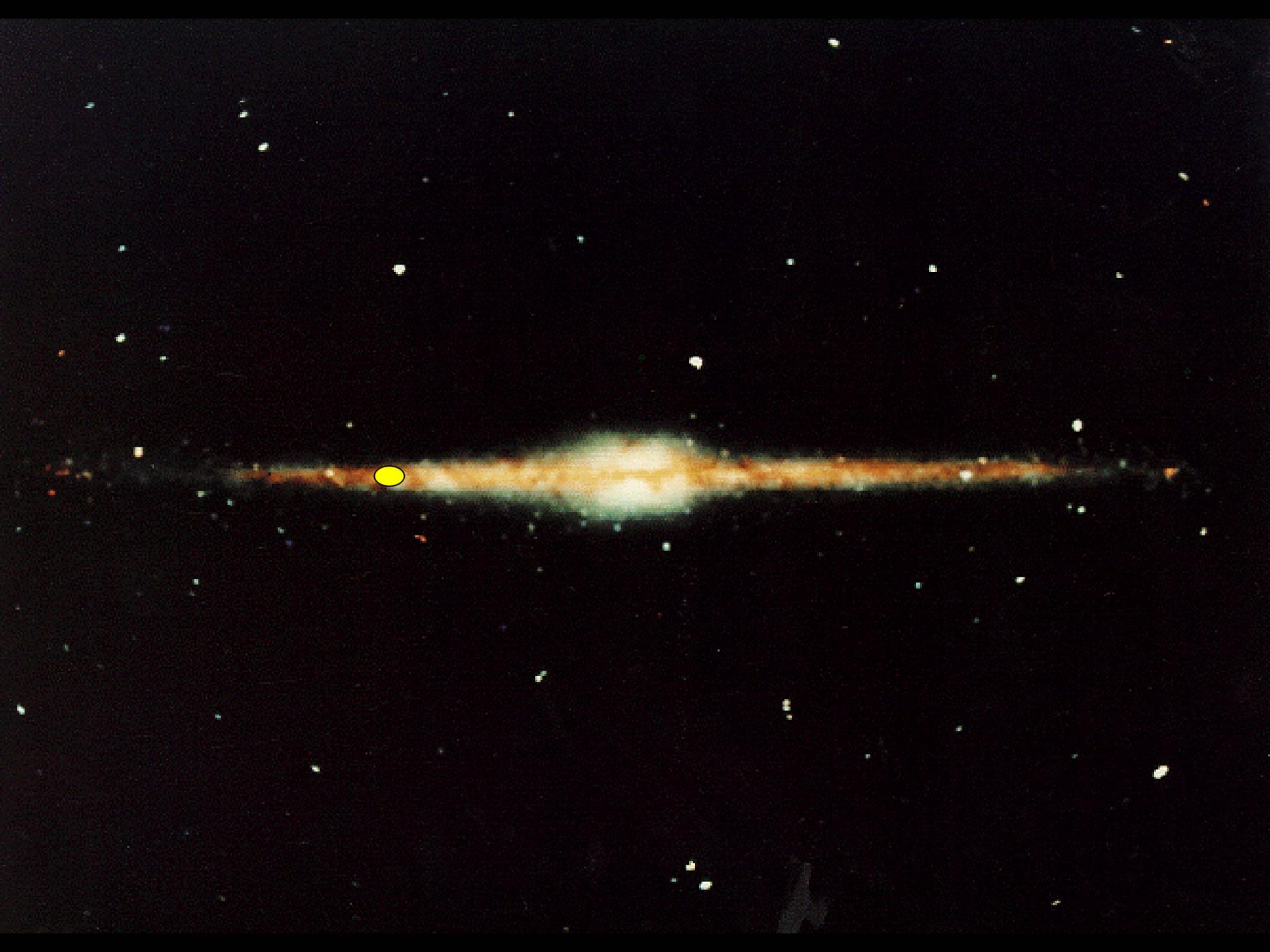


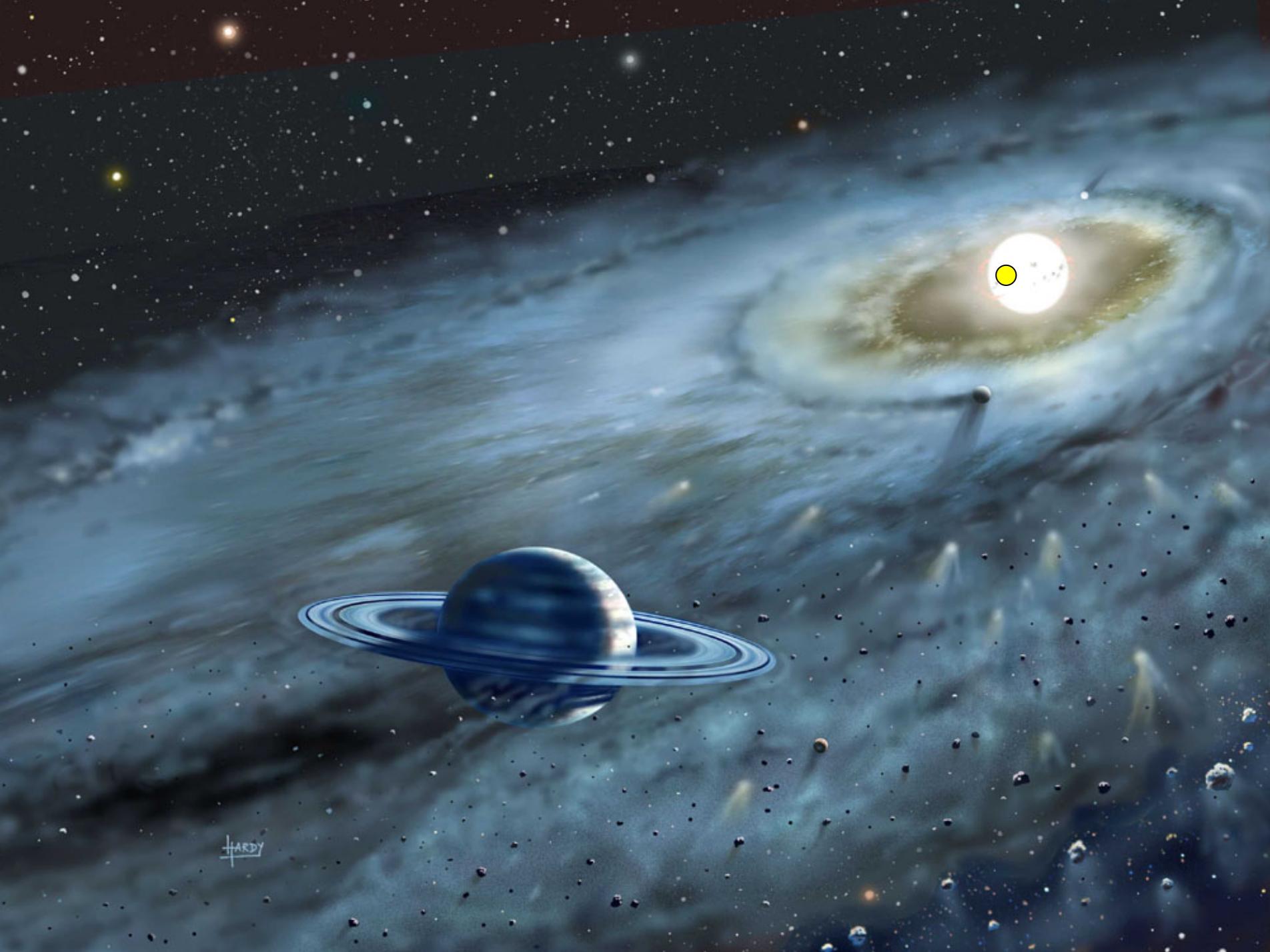
**380,000  
years**

**13.7  
billion  
years**



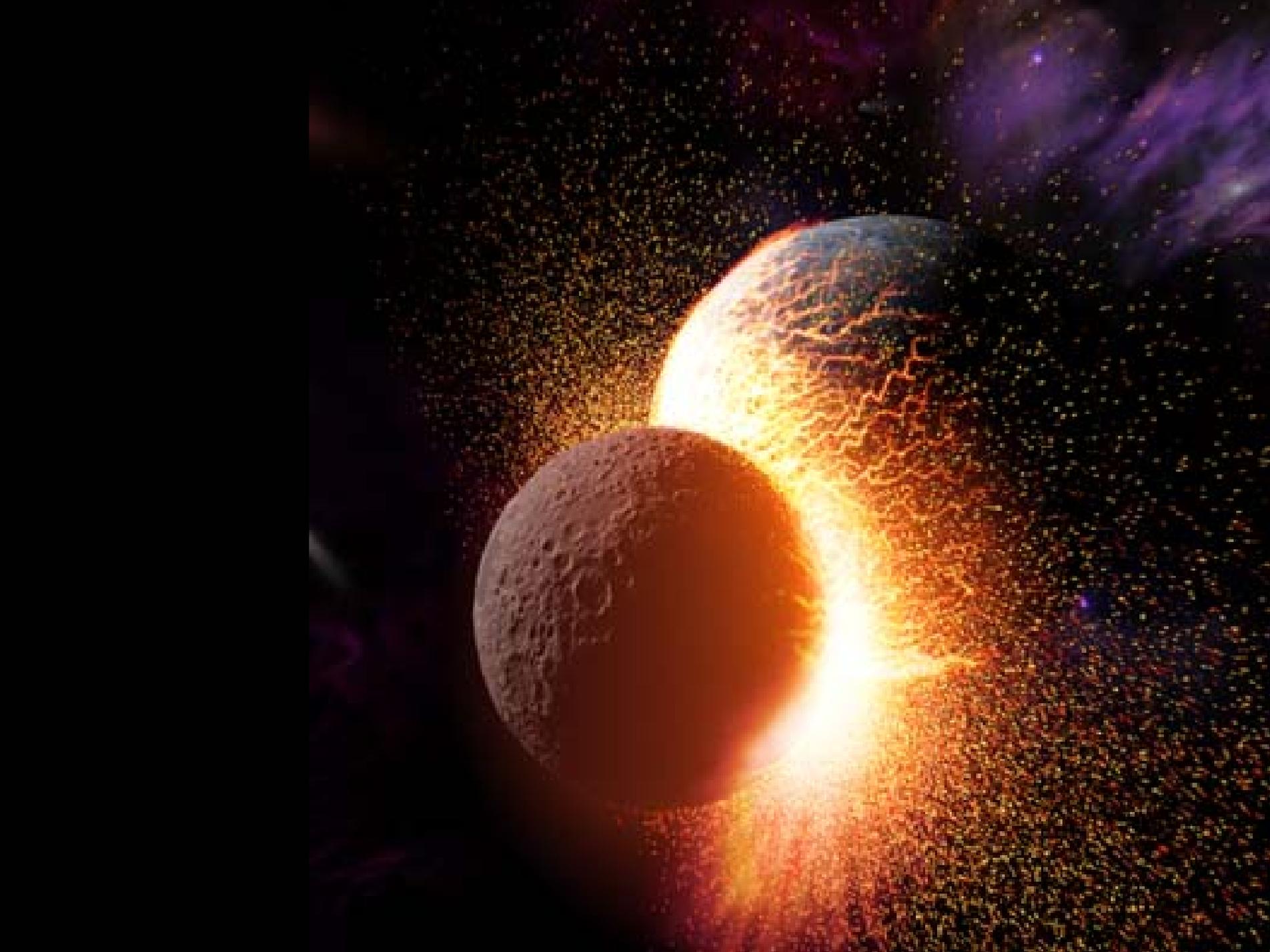






HARDY









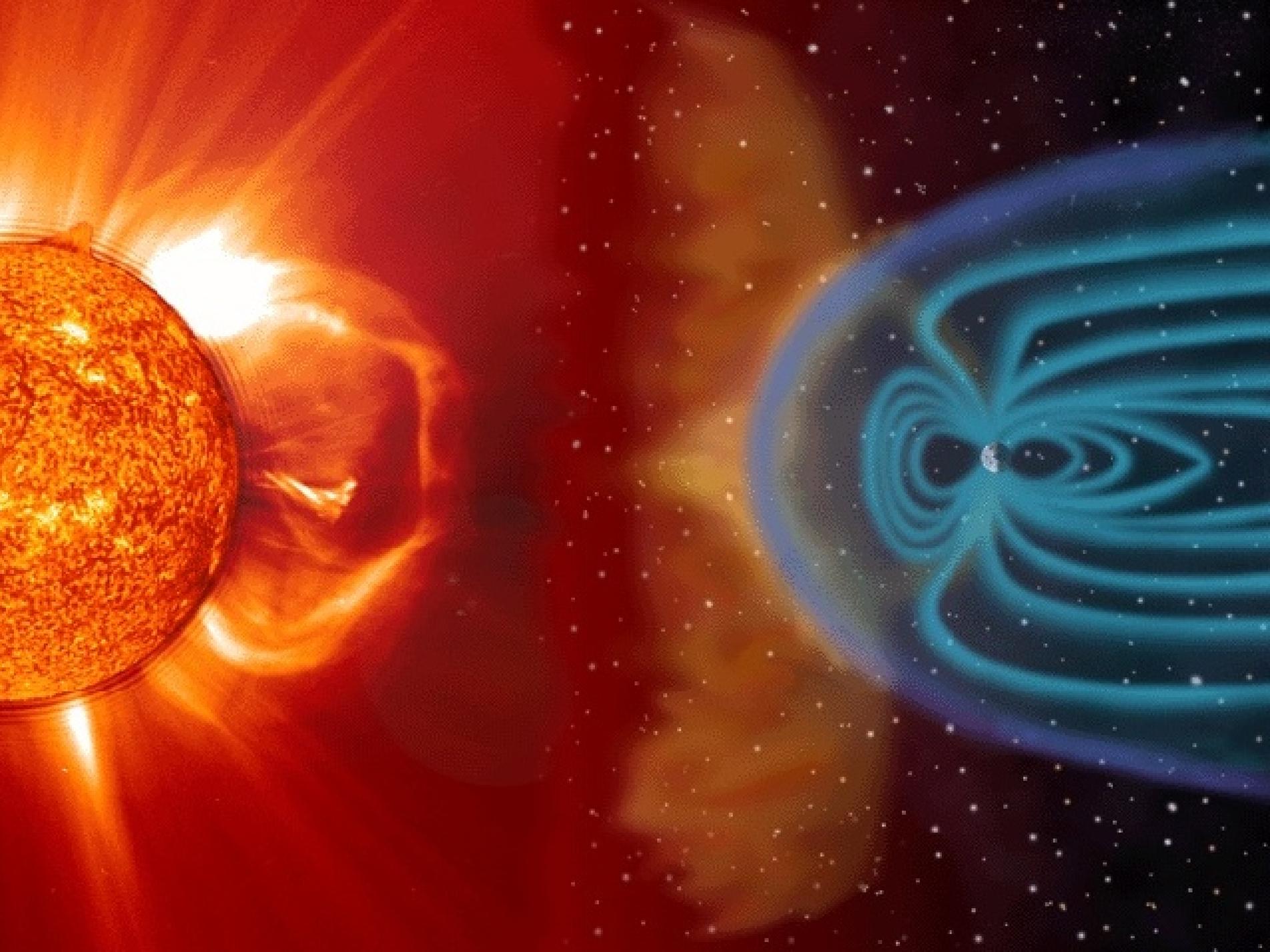
**Planet Earth: an asymmetrical water planet**

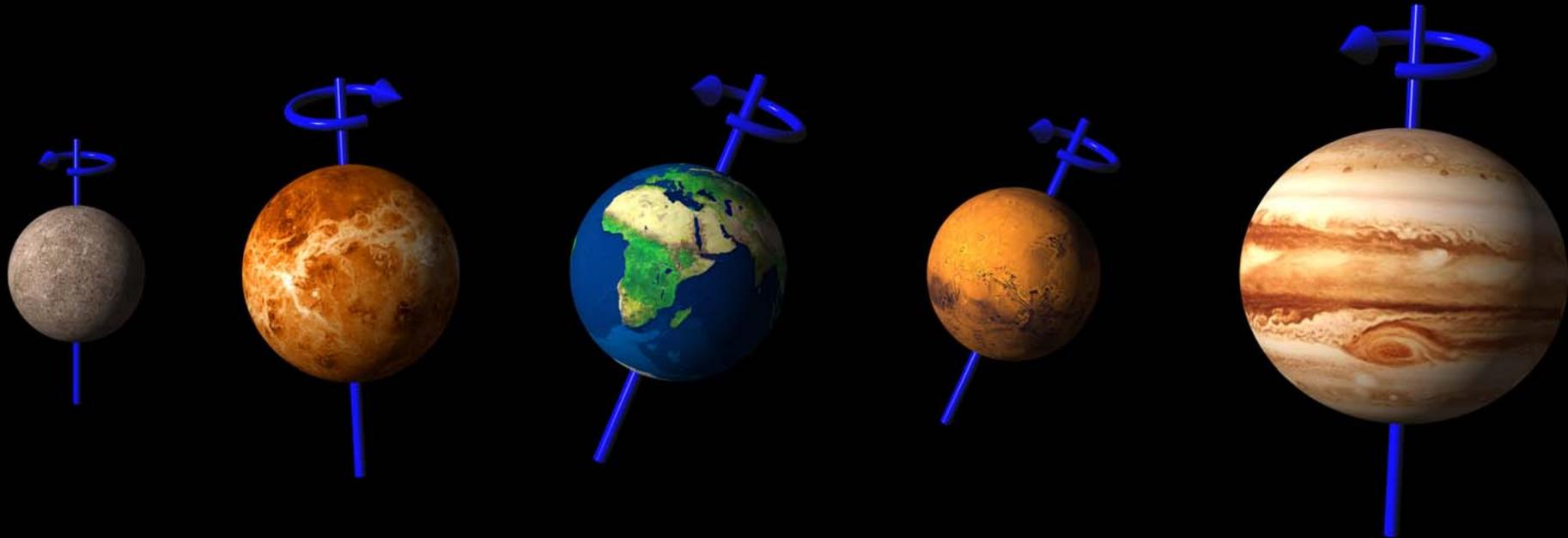
**29% covered by land; 71% covered by oceans**

## Earth Statistics

<b>Mass (kg)</b>	5.976e+24
<b>Mass (Earth = 1)</b>	1.0000e+00
<b>Equatorial radius (km)</b>	6,378.14
<b>Equatorial radius (Earth = 1)</b>	1.0000e+00
<b>Mean density (gm/cm<sup>3</sup>)</b>	5.515
<b>Mean distance from the Sun (km)</b>	149,600,000
<b>Mean distance from the Sun (Earth = 1)</b>	1.0000
<b>Rotational period (days)</b>	0.99727
<b>Rotational period (hours)</b>	23.9345
<b>Orbital period (days)</b>	365.256
<b>Mean orbital velocity (km/sec)</b>	29.79
<b>Orbital eccentricity</b>	0.0167
<b>Tilt of axis (degrees)</b>	23.45
<b>Orbital inclination (degrees)</b>	0.000
<b>Equatorial escape velocity (km/sec)</b>	11.18
<b>Equatorial surface gravity (m/sec<sup>2</sup>)</b>	9.78
<b>Visual geometric albedo</b>	0.37
<b>Mean surface temperature</b>	15°C
<b>Atmospheric pressure (bars)</b>	1.013
<b>Atmospheric composition</b>	
<b>Nitrogen</b>	77%
<b>Oxygen</b>	21%
<b>Other</b>	2%







Mercury  
0.1°

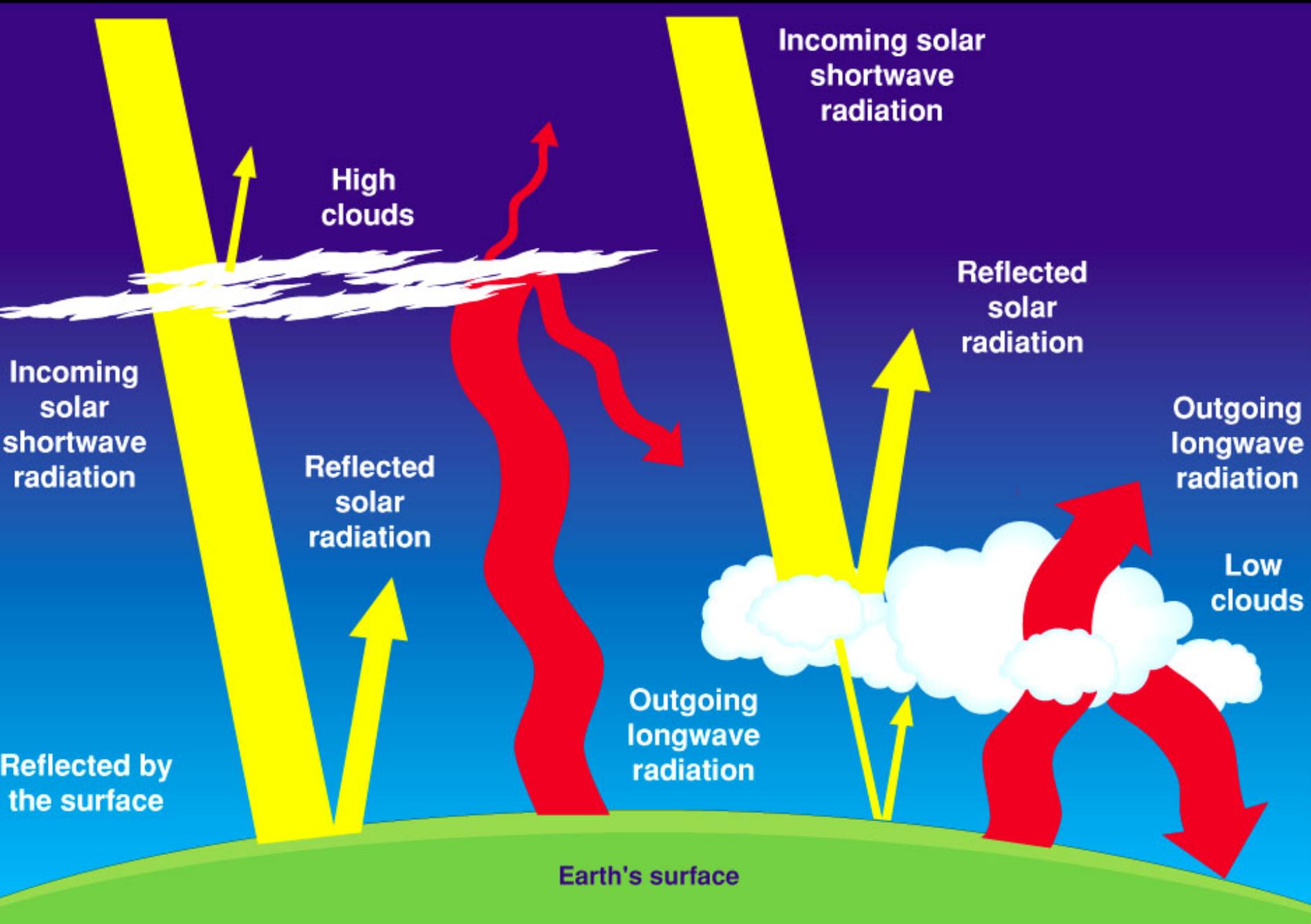
Venus  
177°

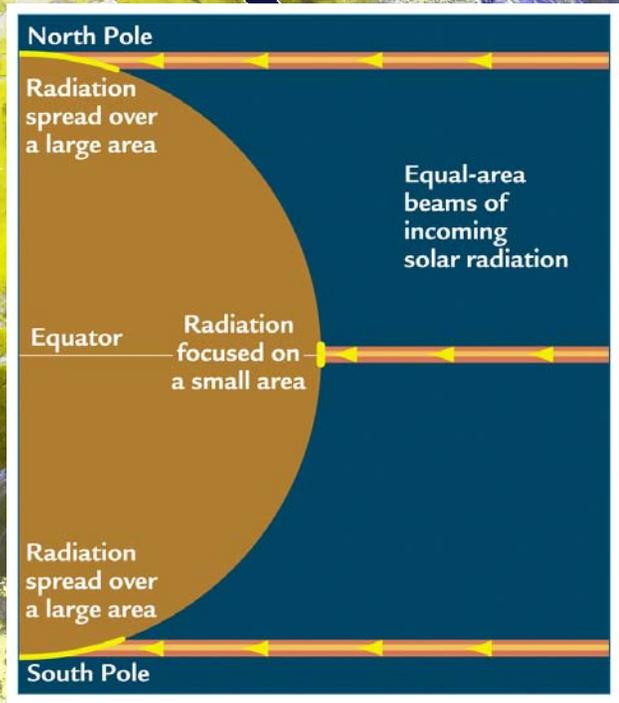
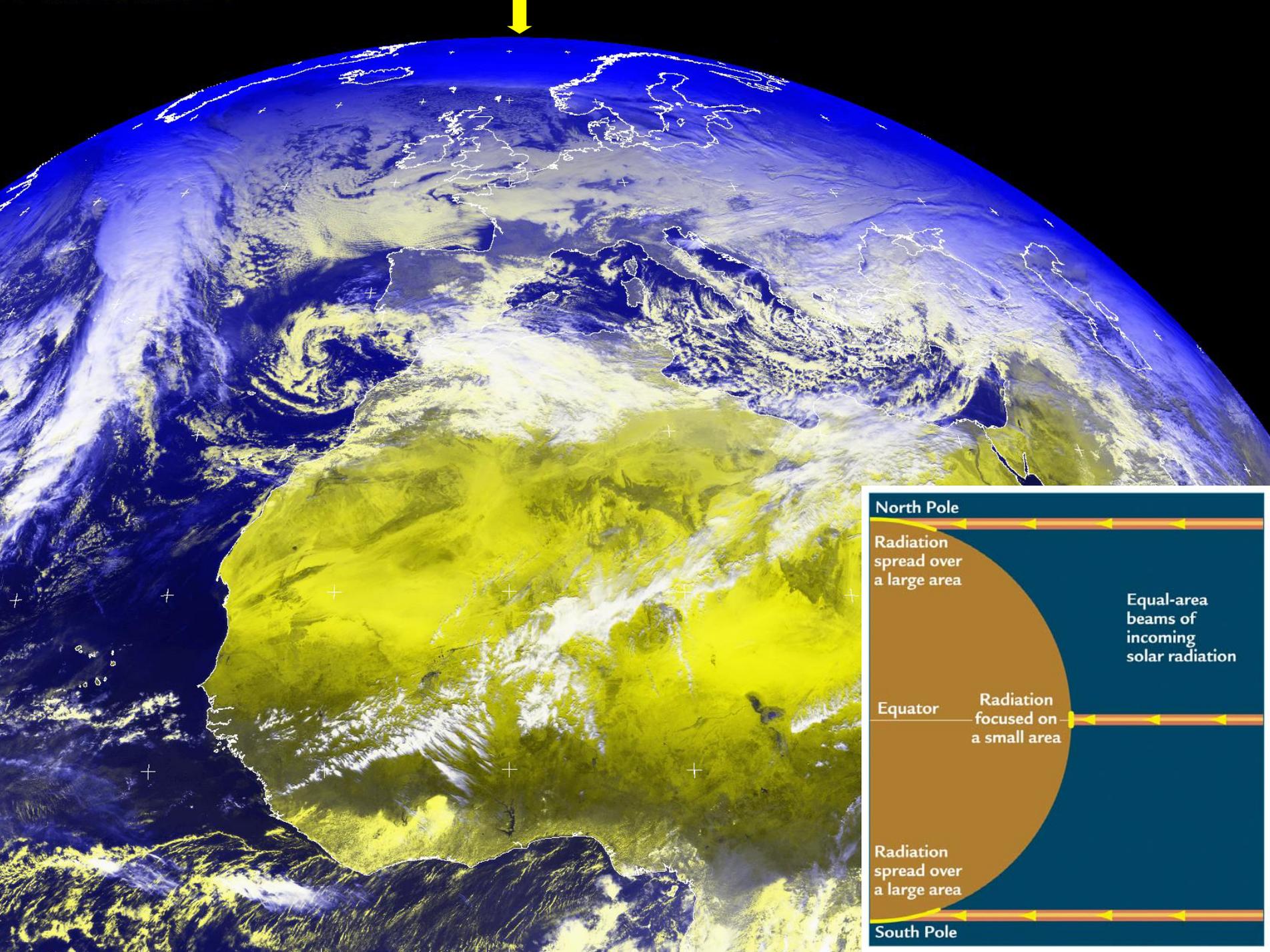
Earth  
23°

Mars  
25°

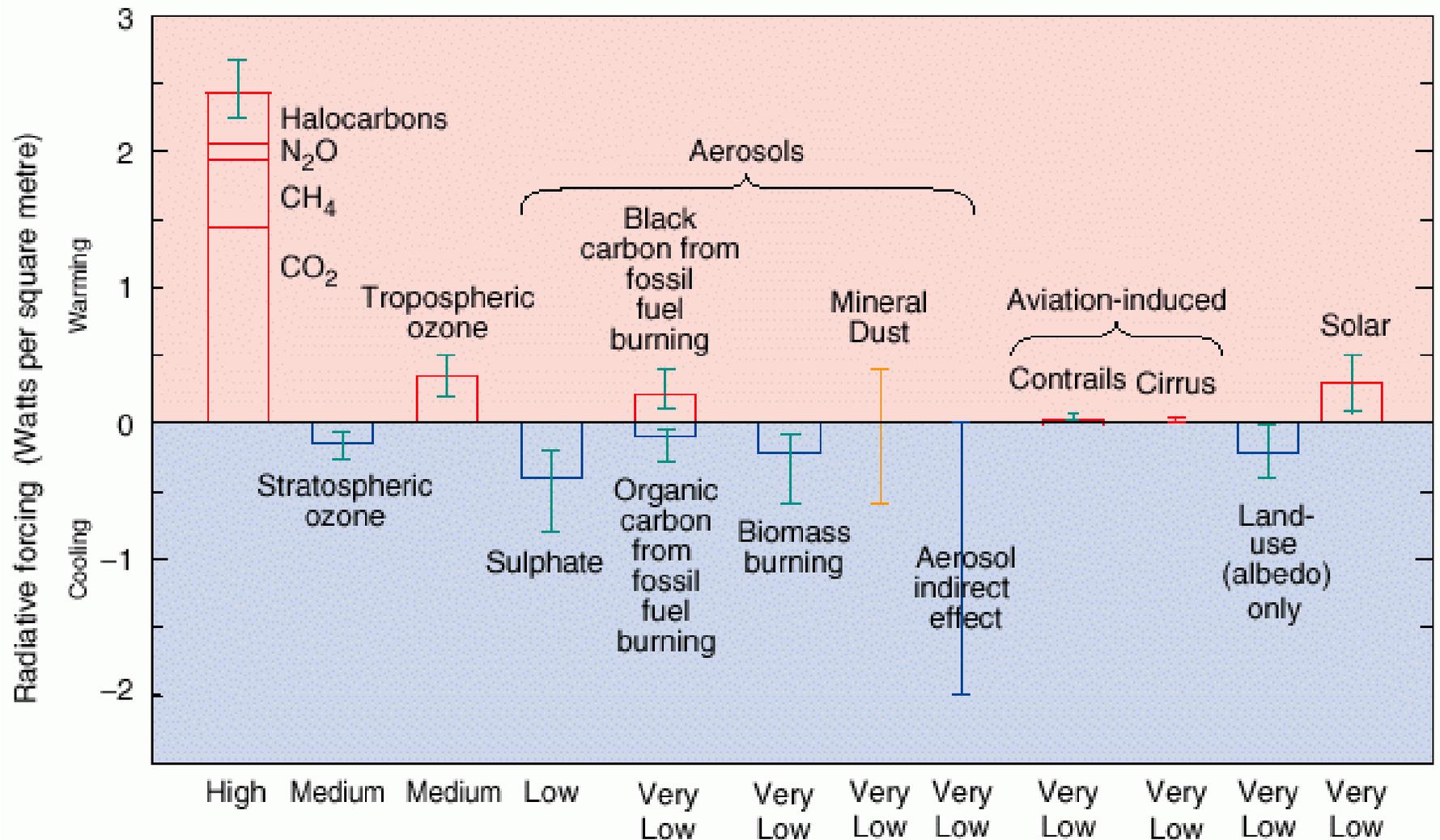
Jupiter  
3°

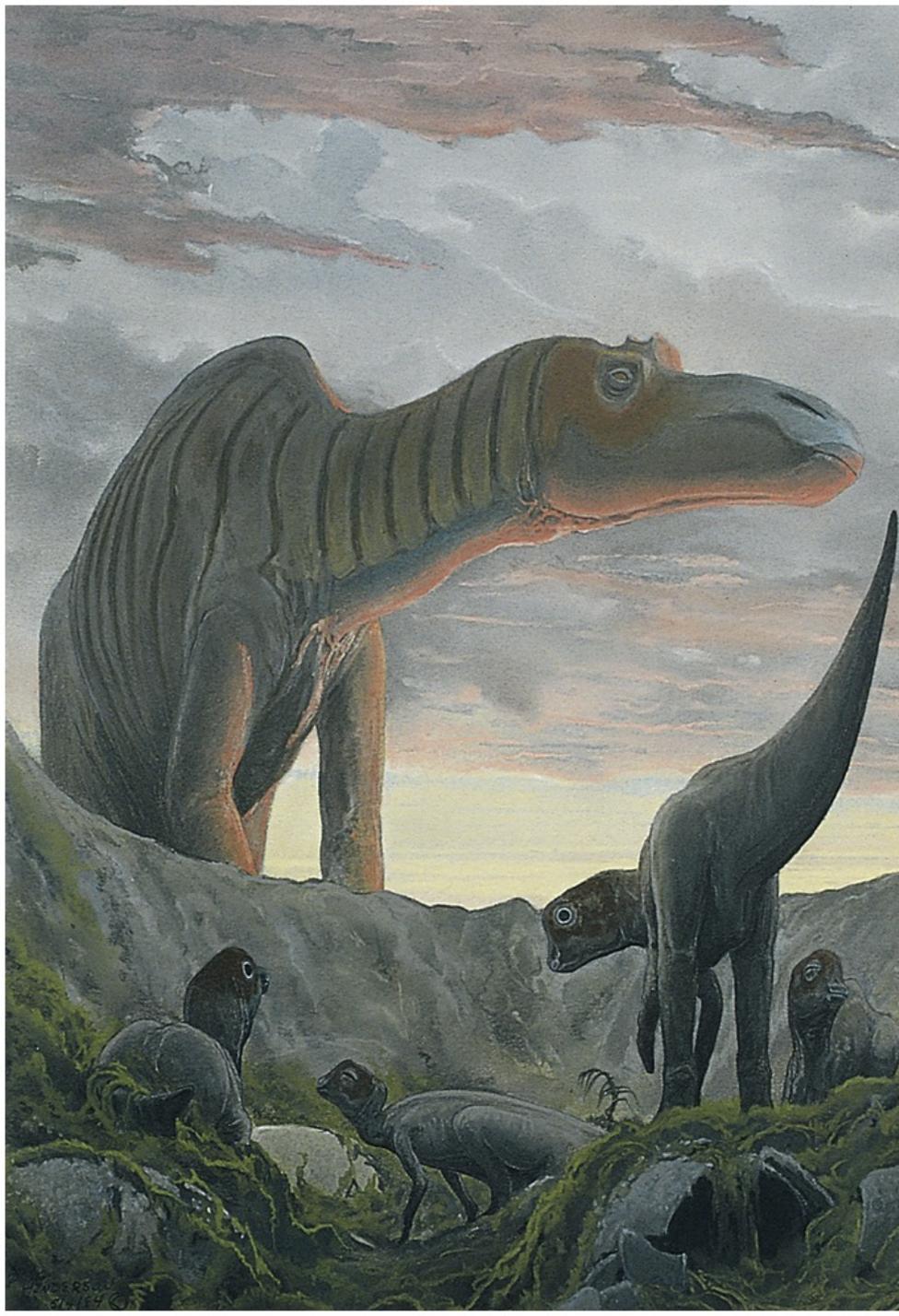
## Obliquity of the Nine Planets





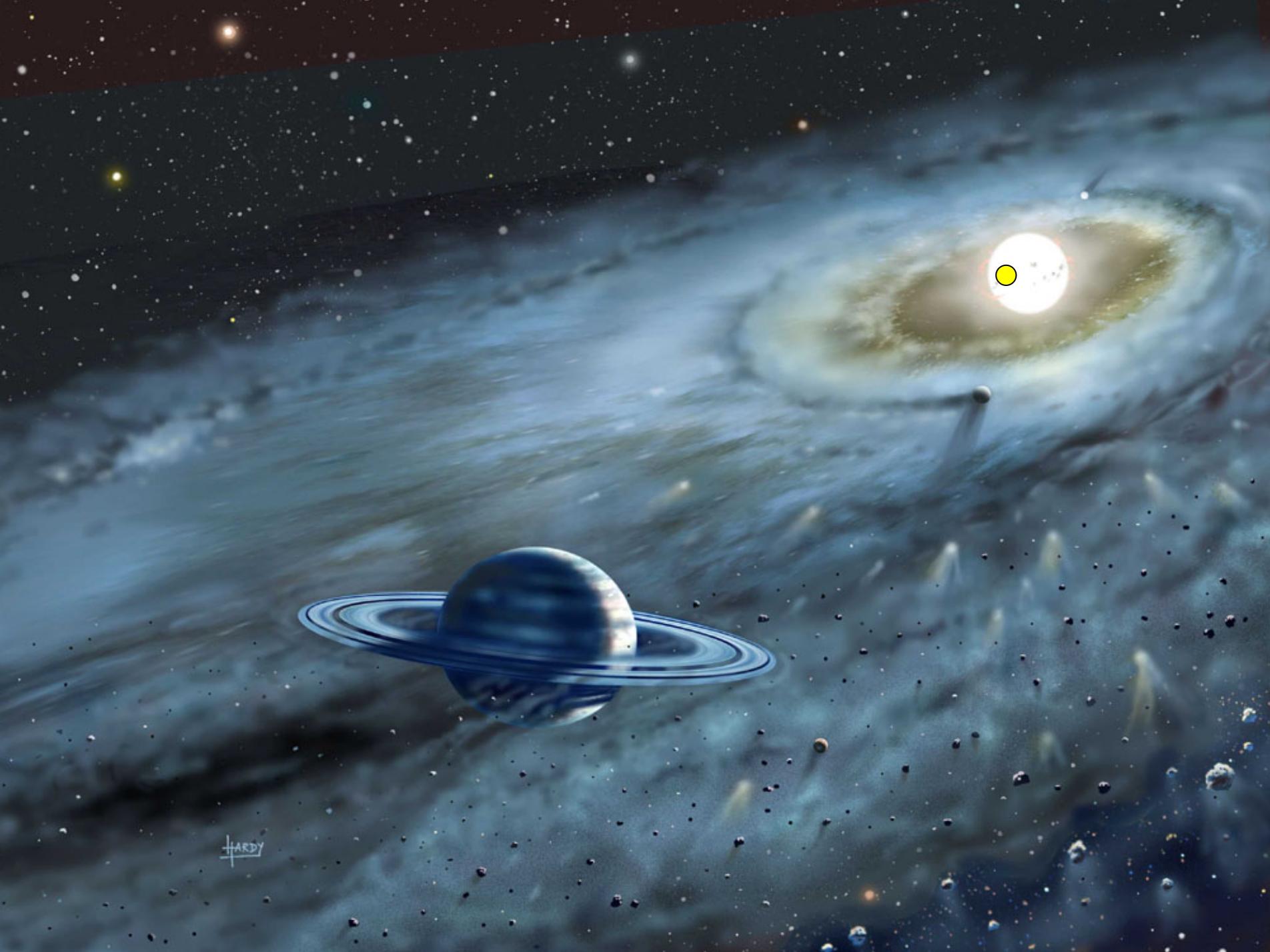
# The global mean radiative forcing of the climate system for the year 2000, relative to 1750



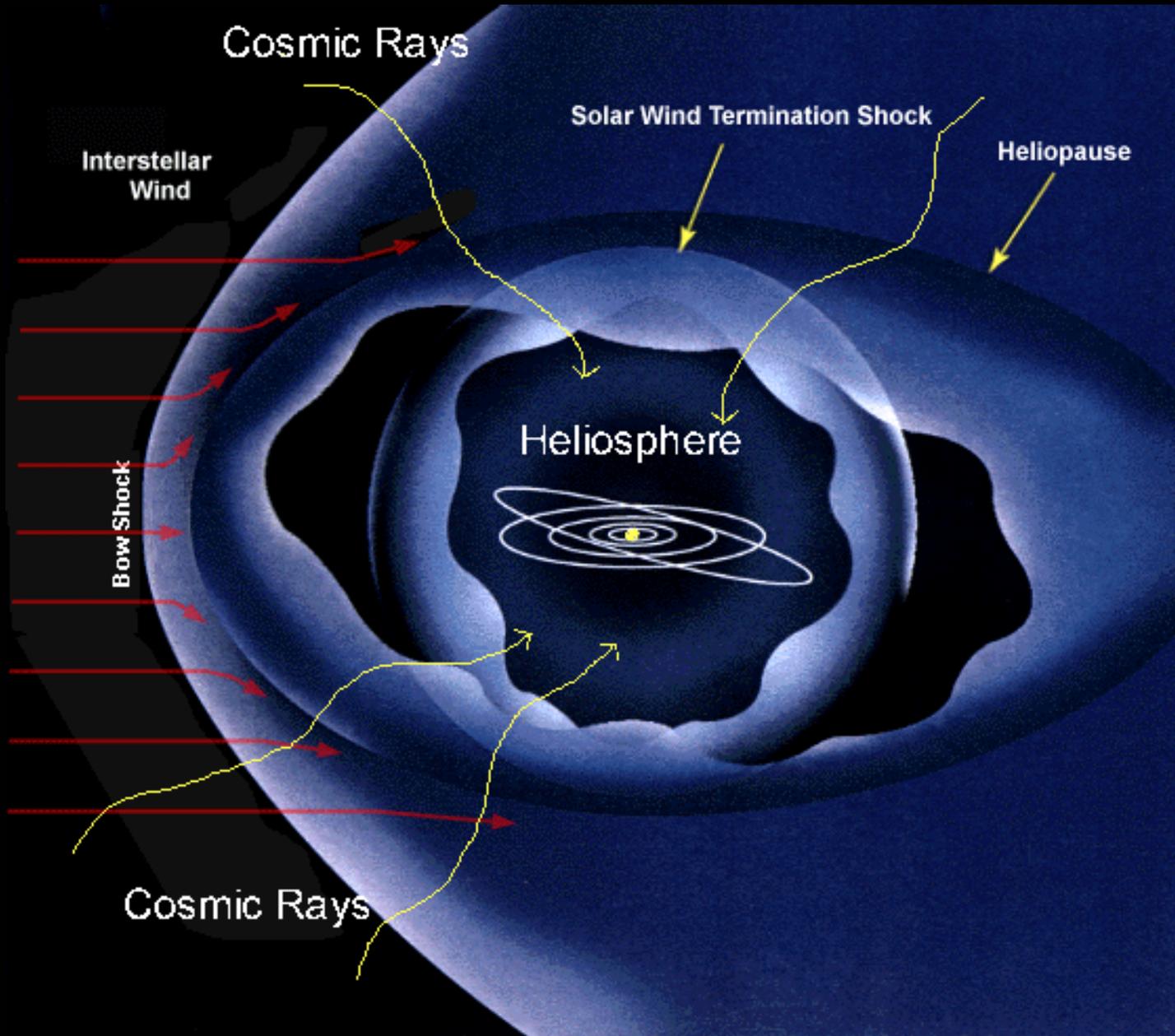




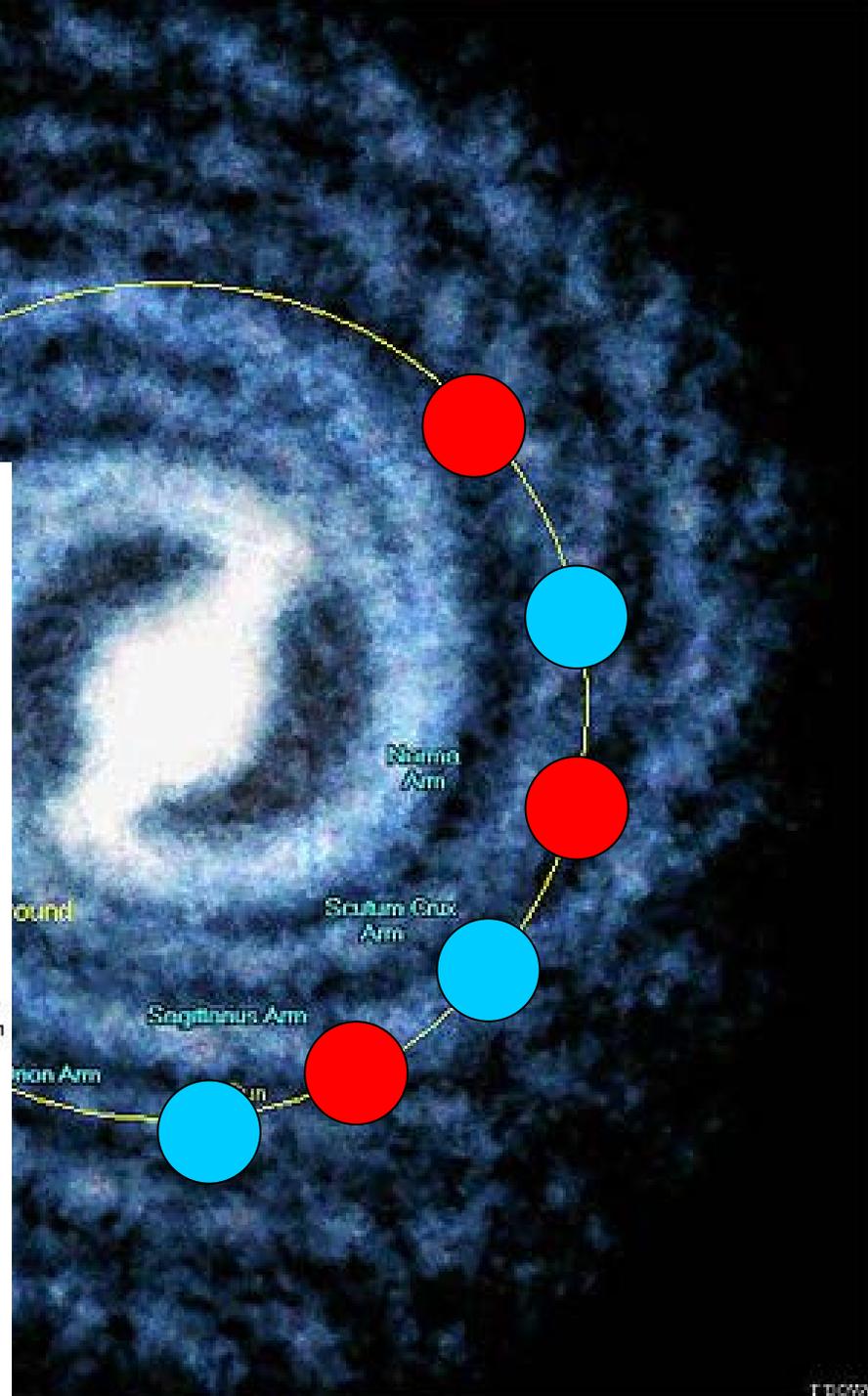
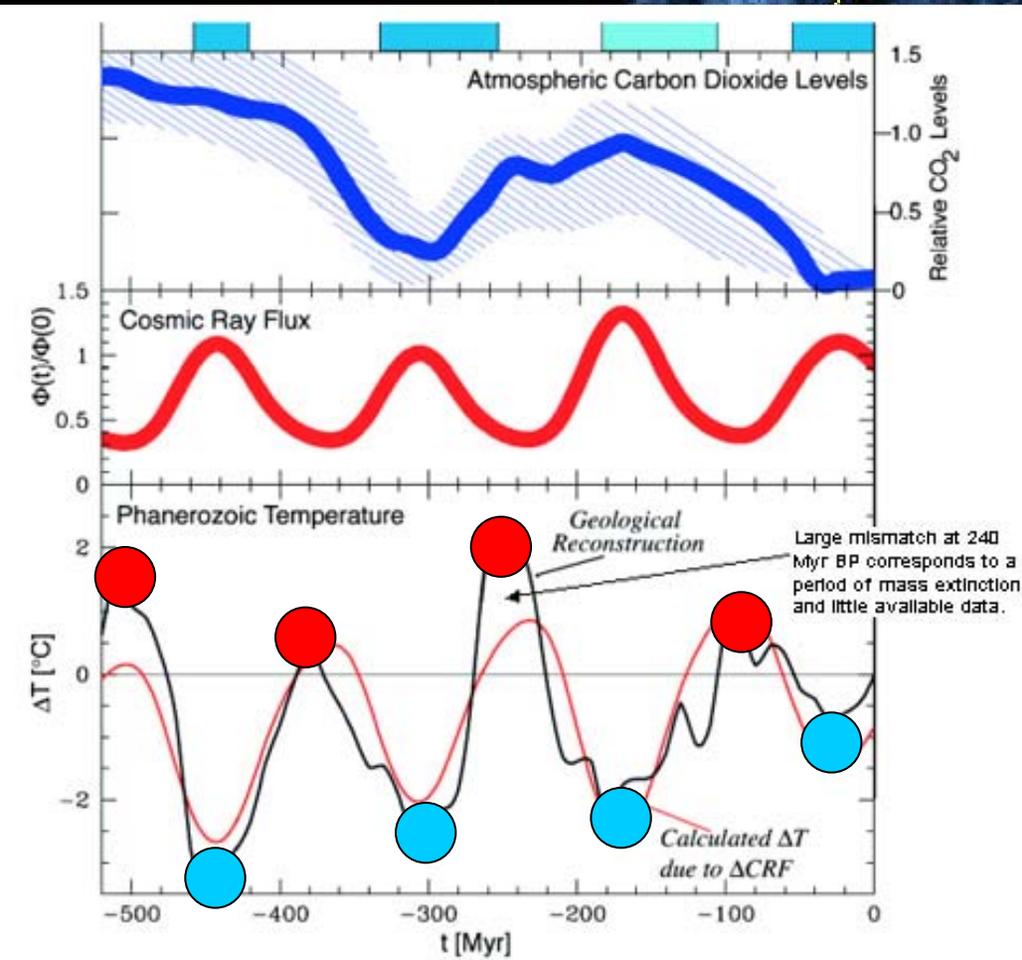
This is it, Fredman. Unmatched evidence of a surprisingly rapid climatic change at the onset of the last ice age !



HARDY

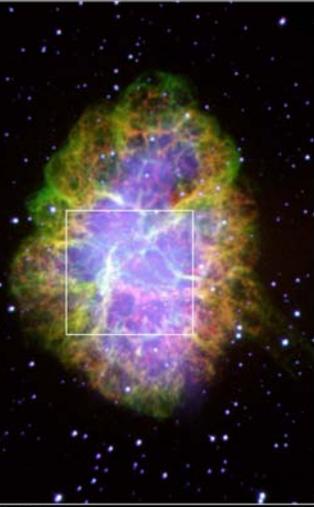




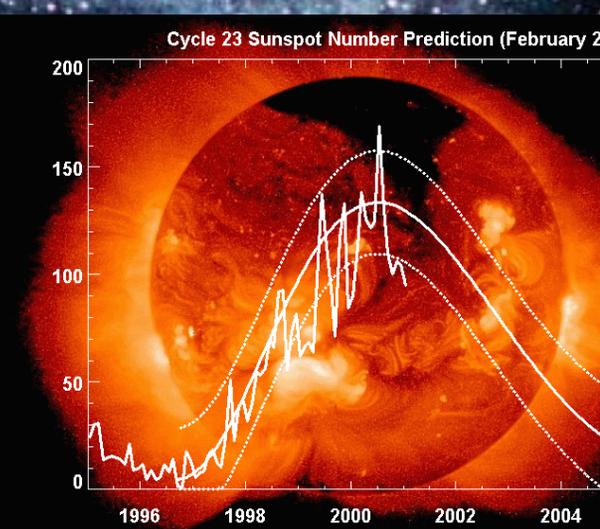
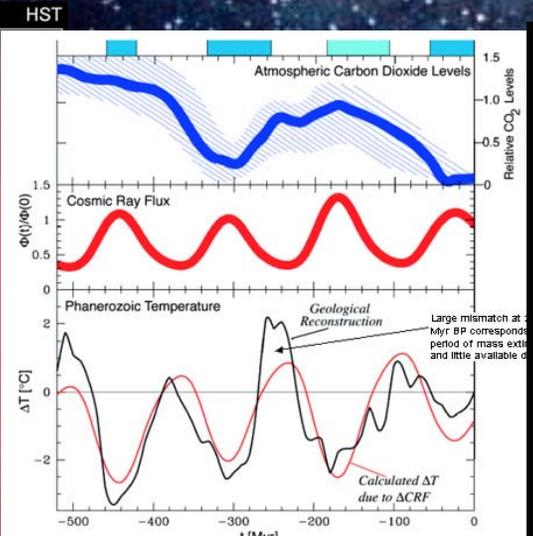
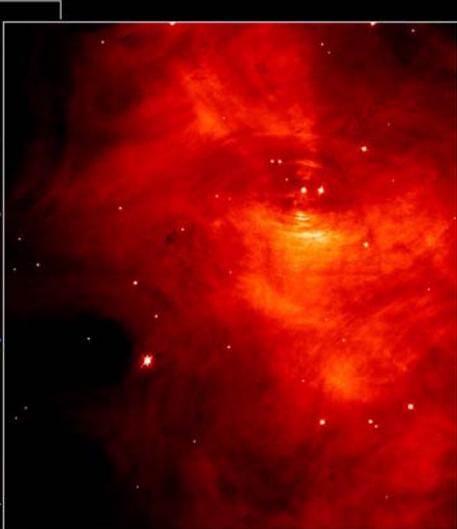




Supernova  
Crab Nebula observed 1045 AD



Palomar



**1: Variations in cosmic radiation may cause climatic changes**

**2: Variations in solar activity may cause climatic changes**



## **Milutin Milankovitch 1879-1958**

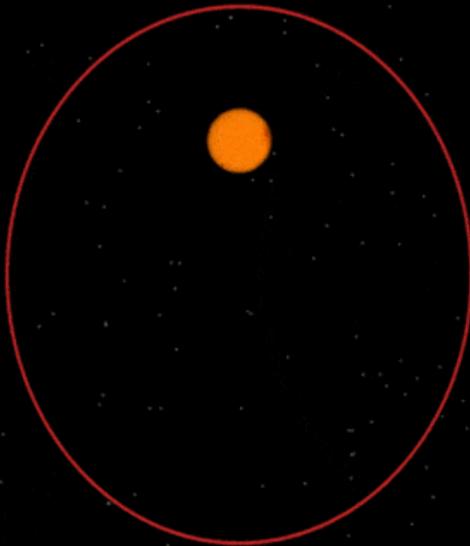
The Serbian astrophysicist Milutin Milankovitch dedicated his career to developing a mathematical theory of climate based on the seasonal and latitudinal variations of solar radiation received by the Earth.

Now known as the Milankovitch Theory, it states that as the Earth travels through space around the sun, cyclical variations in three elements of Earth-sun geometry combine to produce variations in the amount of solar energy that reaches Earth.

### Variation in Orbital Eccentricity



eccentricity = 0



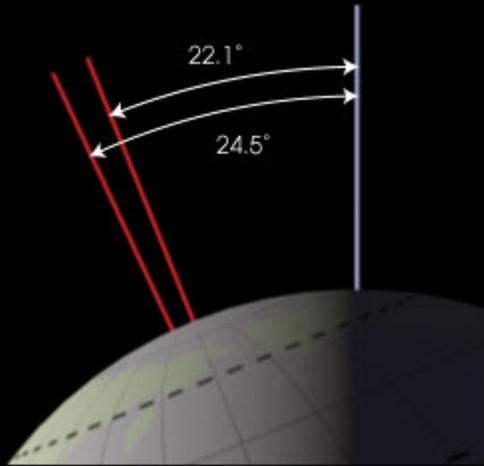
eccentricity = .5

Changes in **orbital eccentricity** affect the Earth-sun distance.

Currently, a difference of only 3 percent (5 million kilometers) exists between closest approach (perihelion), which occurs on or about January 3, and furthest departure (aphelion), which occurs on or about July 4. This difference in distance amounts to about a 6 percent increase in incoming solar radiation (insolation) from July to January.

The shape of the Earth's orbit changes from being elliptical (high eccentricity) to being nearly circular (low eccentricity) in a cycle that takes between 90,000 and 100,000 years. When the orbit is highly elliptical, the amount of insolation received at perihelion would be on the order of 20 to 30 percent greater than at aphelion, resulting in a substantially different climate from what we experience today.

Variation in Axial Obliquity



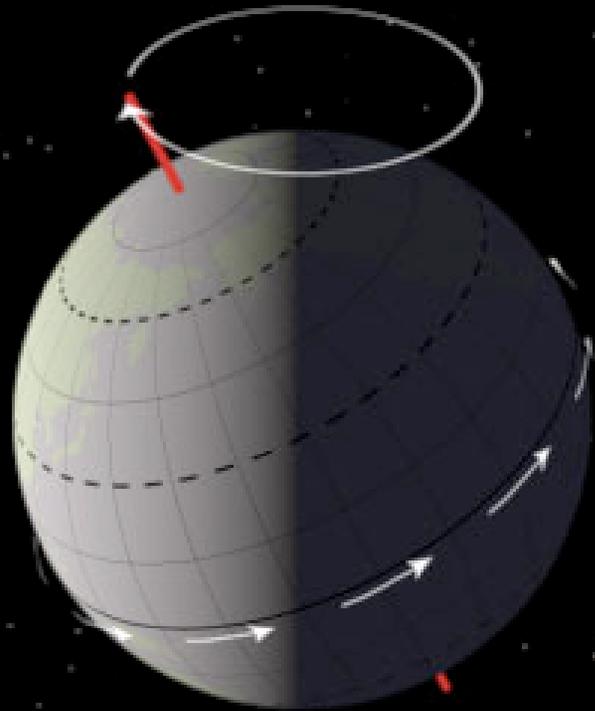
## Obliquity (change in axial tilt)

As the axial tilt increases, the seasonal contrast increases so that winters are colder and summers are warmer in both hemispheres.

Today, the Earth's axis is tilted 23.5 degrees from the plane of its orbit around the sun. But this tilt changes. During a cycle that averages about 40,000 years, the tilt of the axis varies between 22.1 and 24.5 degrees.

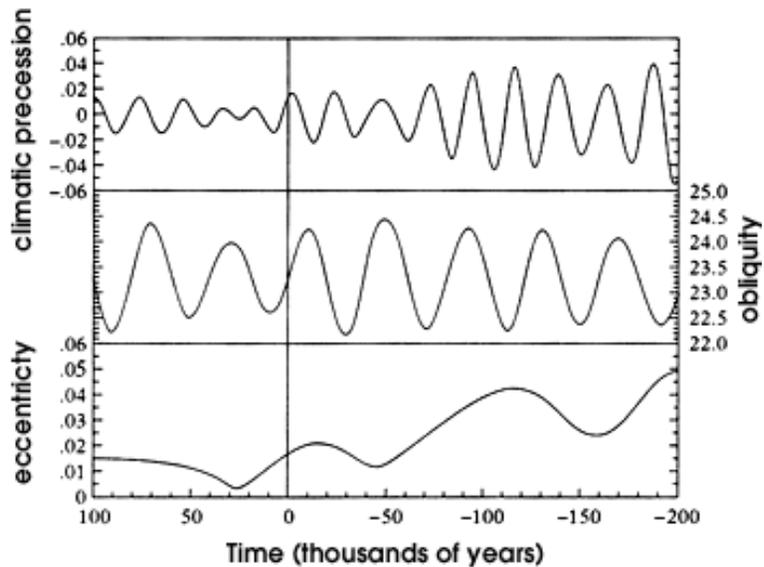
More tilt means more severe seasons—warmer summers and colder winters; less tilt means less severe seasons—cooler summers and milder winters. It's the cool summers that are thought to allow snow and ice to last from year-to-year in high latitudes, eventually building up into massive ice sheets.

## Precession



## Precession

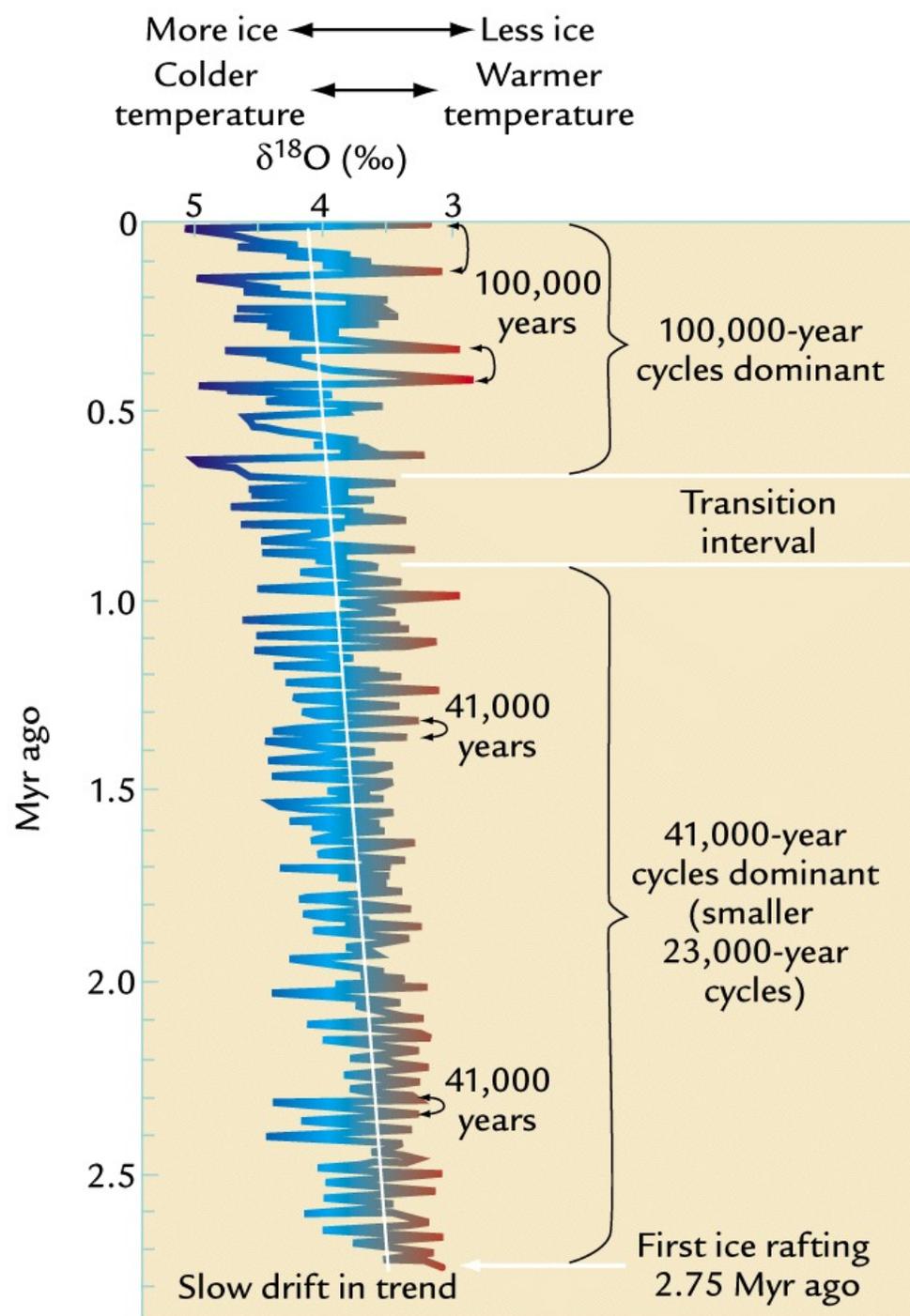
Changes in axial precession alter the dates of perihelion and aphelion, and therefore increase the seasonal contrast in one hemisphere and decrease the seasonal contrast in the other hemisphere.

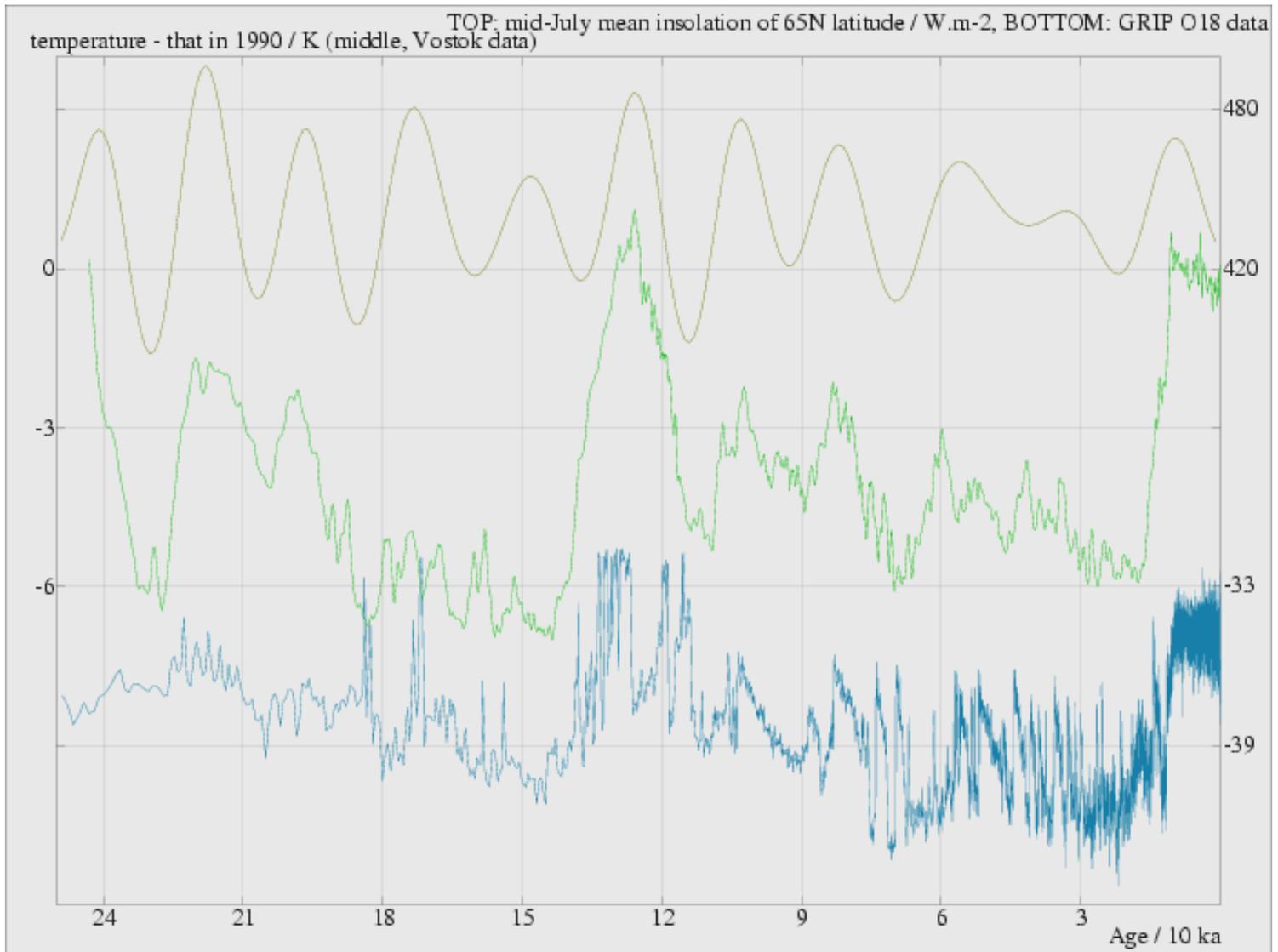


Graph showing the calculated values for 300,000 years of orbital variation by Berger and Loutre, 1991.

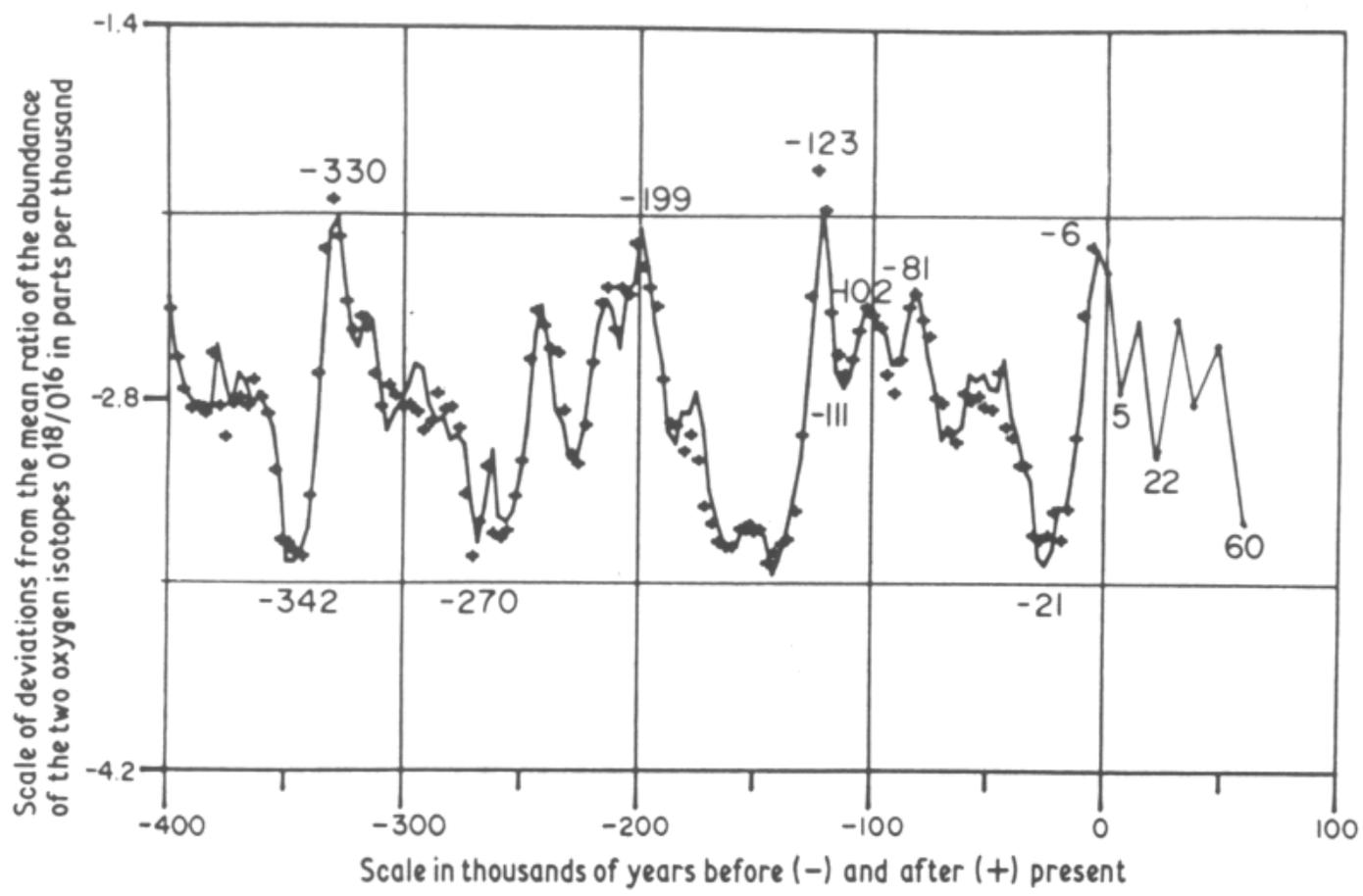
For about 50 years, however, Milankovitch's theory was largely ignored.

Using these three orbital variations, Milankovitch was able to formulate a comprehensive mathematical model that calculated latitudinal differences in insolation and the corresponding surface temperature for 600,000 years prior to the year 1800. He then attempted to correlate these changes with the growth and retreat of the Ice Ages. To do this, Milankovitch assumed that radiation changes in some latitudes and seasons are more important to ice sheet growth and decay than those in others. Then, at the suggestion of German Climatologist Vladimir Köppen, he chose summer insolation at 65 degrees North as the most important latitude and season to model, reasoning that great ice sheets grew near this latitude and that cooler summers might reduce summer snowmelt, leading to a positive annual snow budget and ice sheet growth.

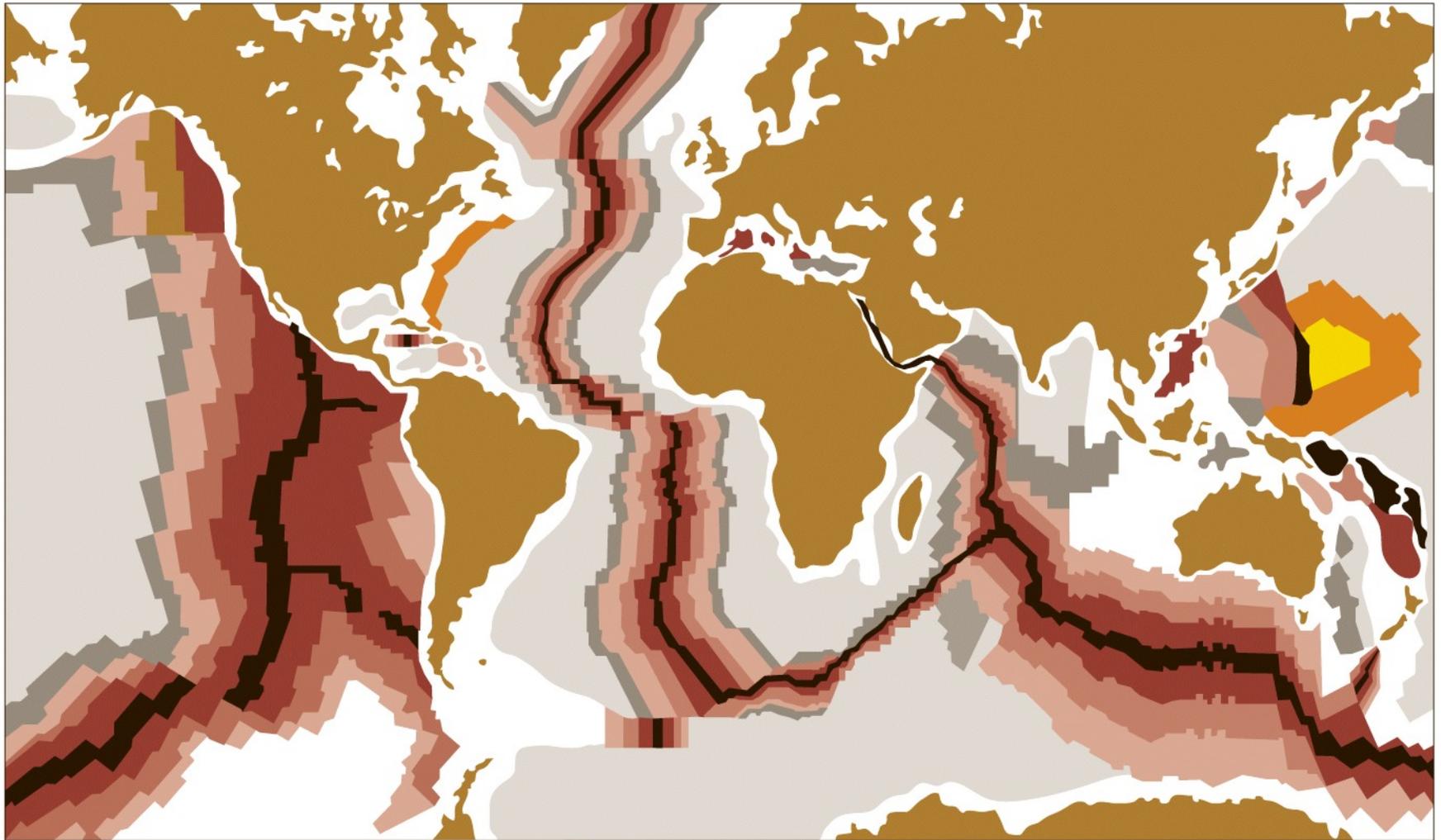




Climate and insolation since 240.000 BP.

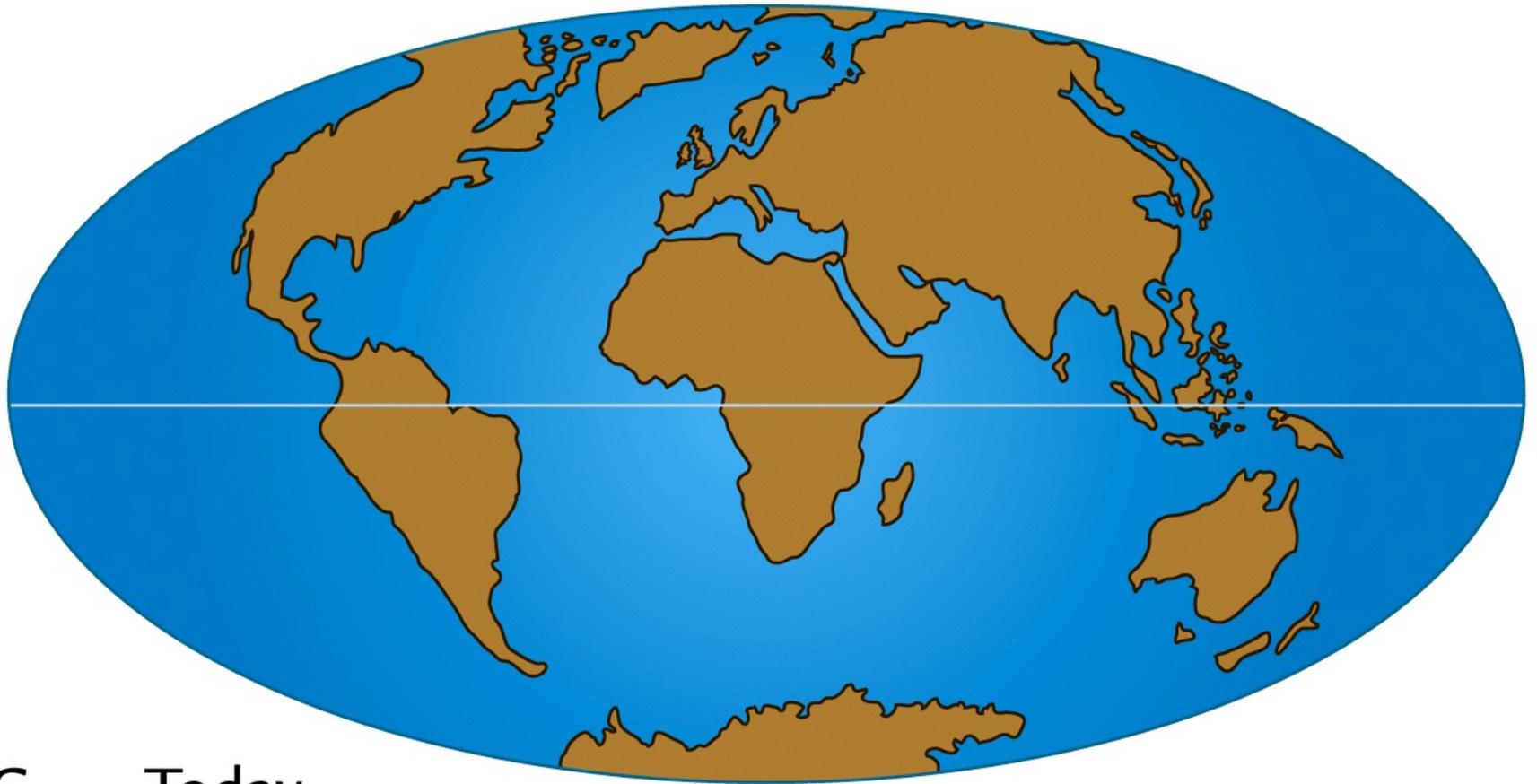


**Orbital orbital variations may cause climatic changes**

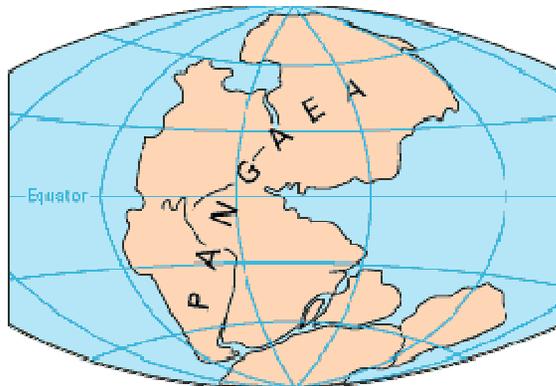


Age (Myr)

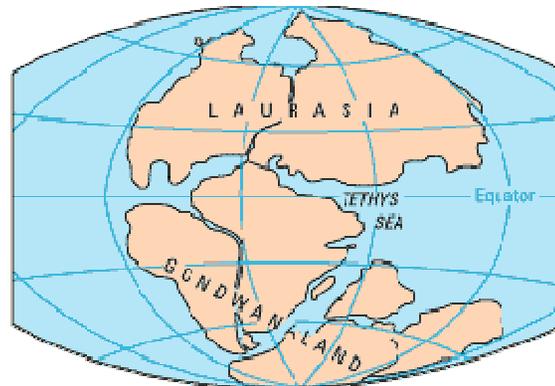
0 - 5	5 - 21	21 - 38	38 - 52
52 - 65	65 - 140	140 - 160	> 160



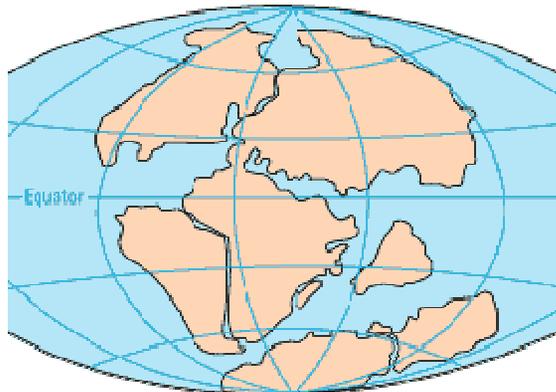
C Today



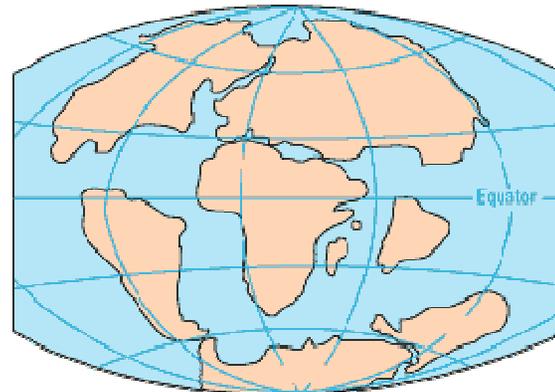
**PERMIAN**  
225 million years ago



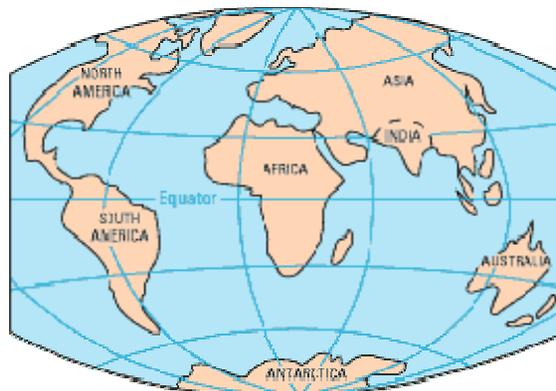
**TRIASSIC**  
200 million years ago



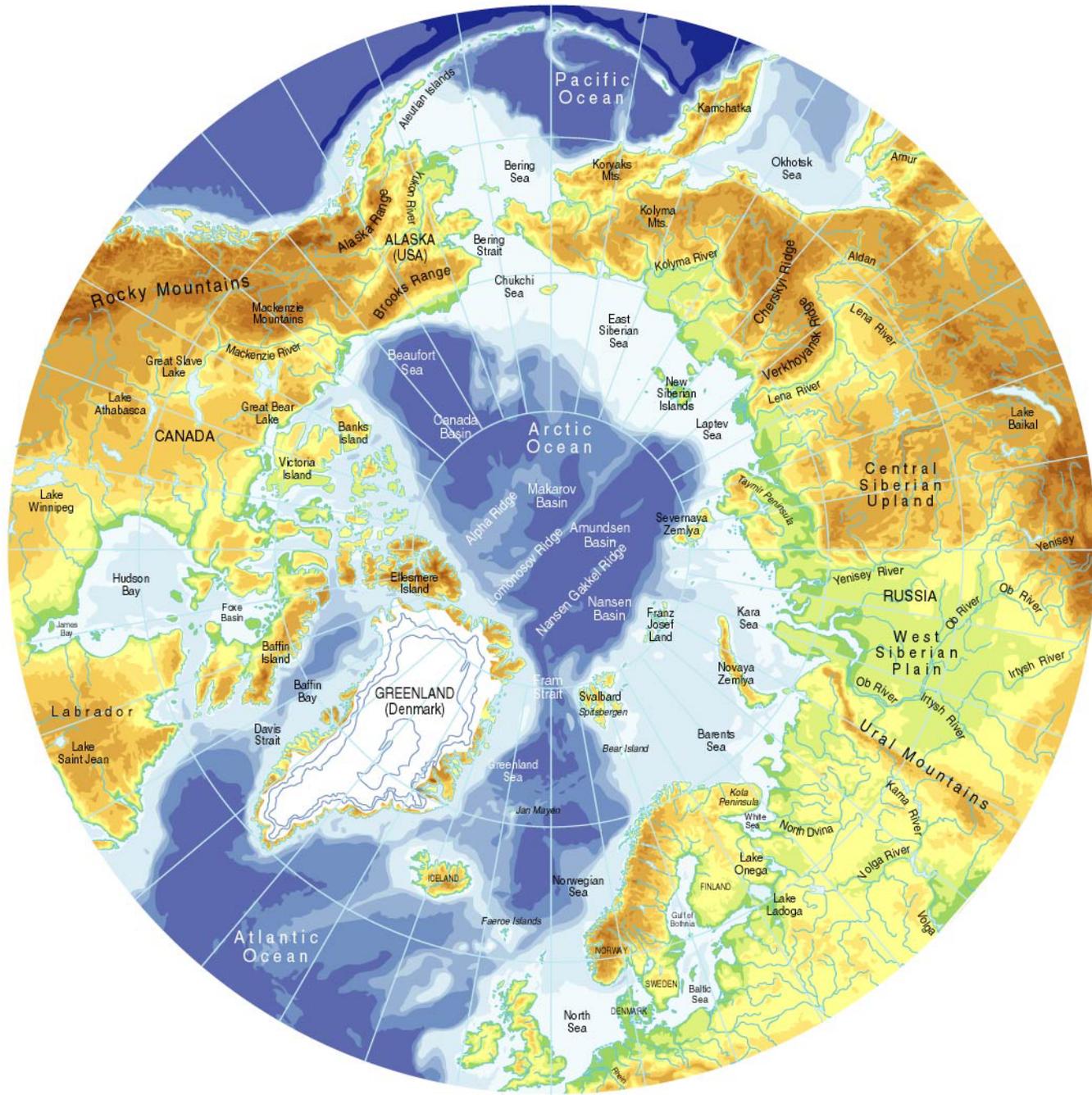
**JURASSIC**  
135 million years ago



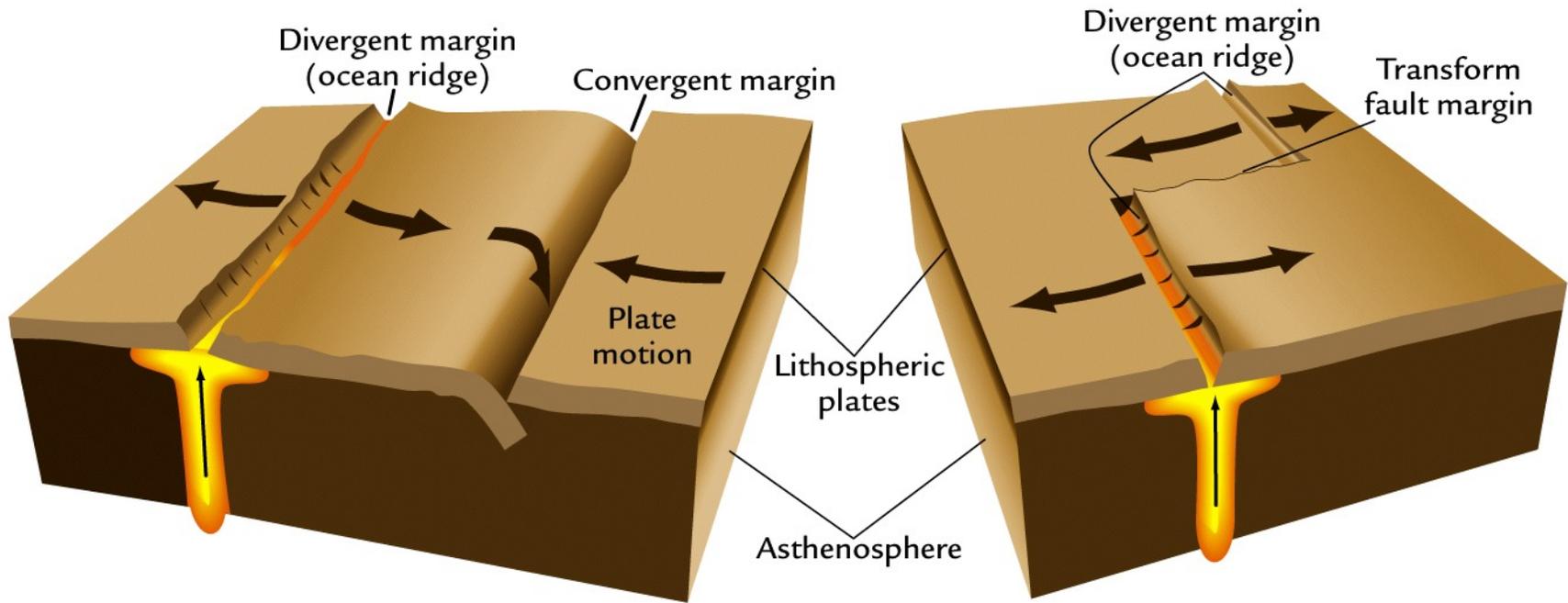
**CRETACEOUS**  
65 million years ago

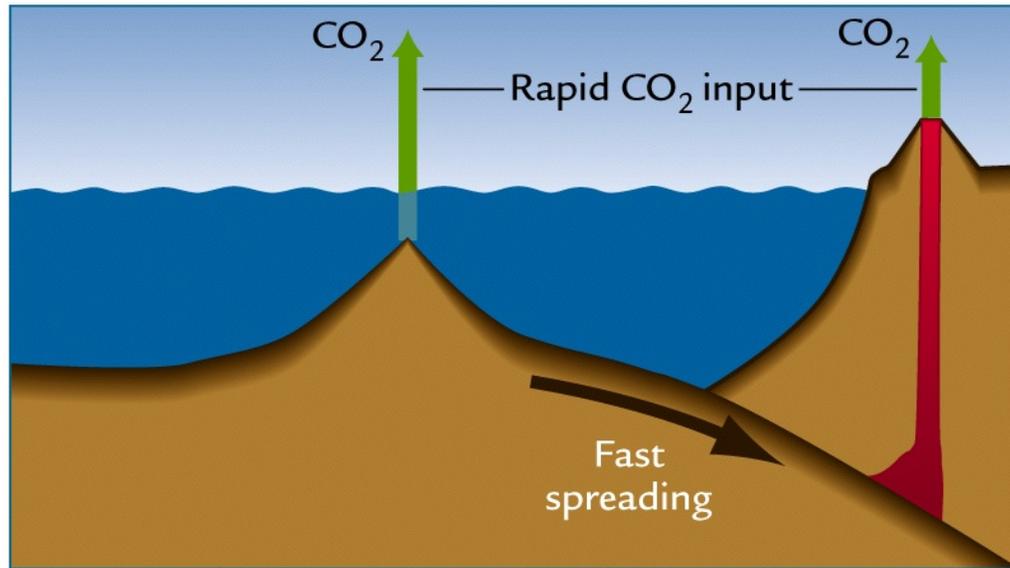


**PRESENT DAY**

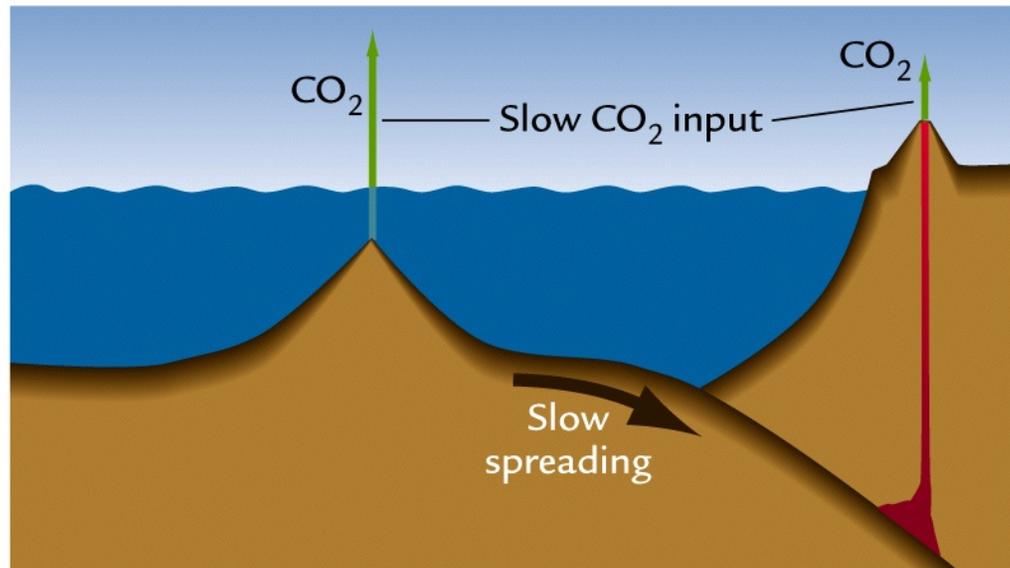


**Continental drift may cause climatic changes by changing the land-sea configuration**

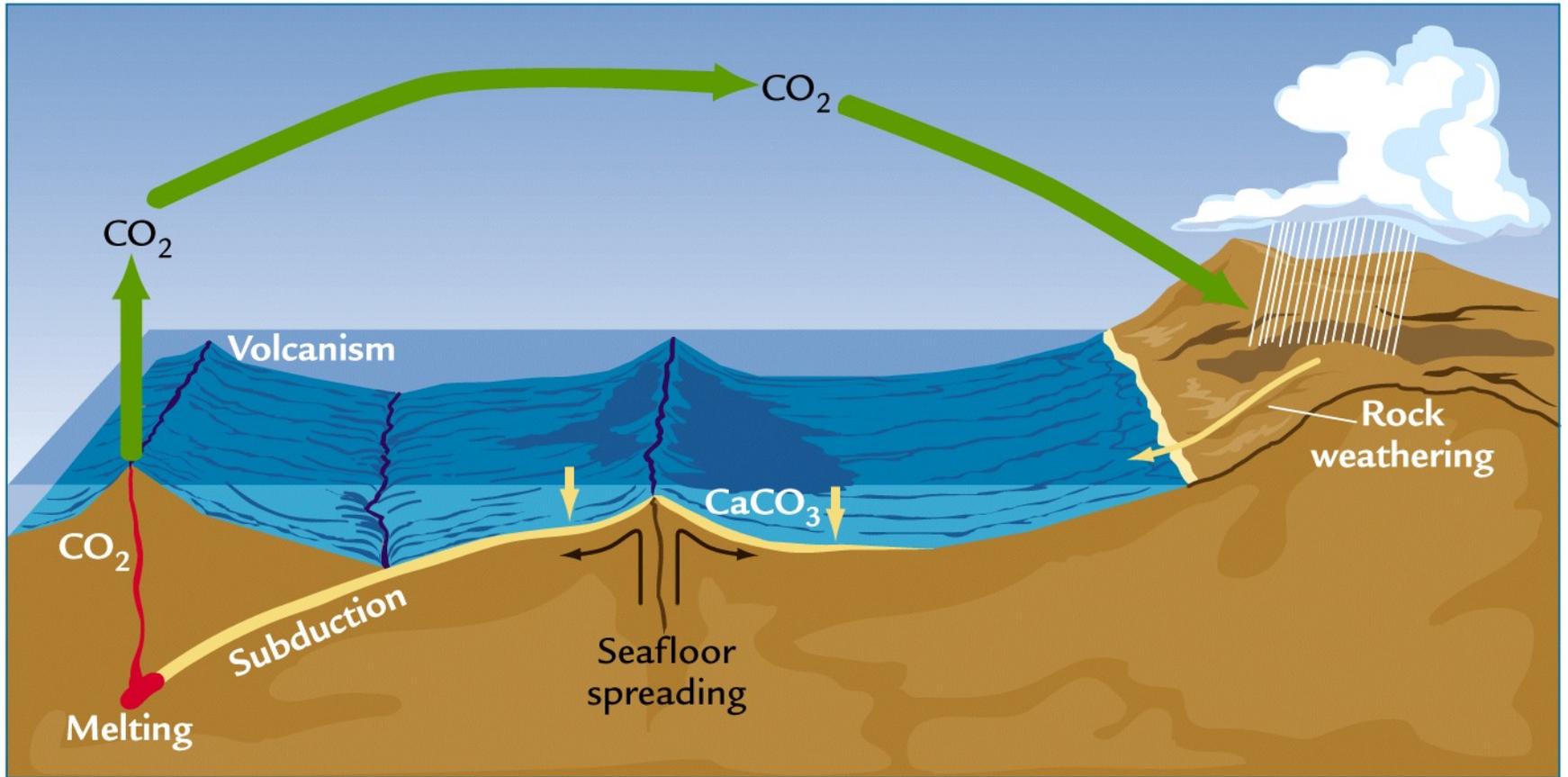


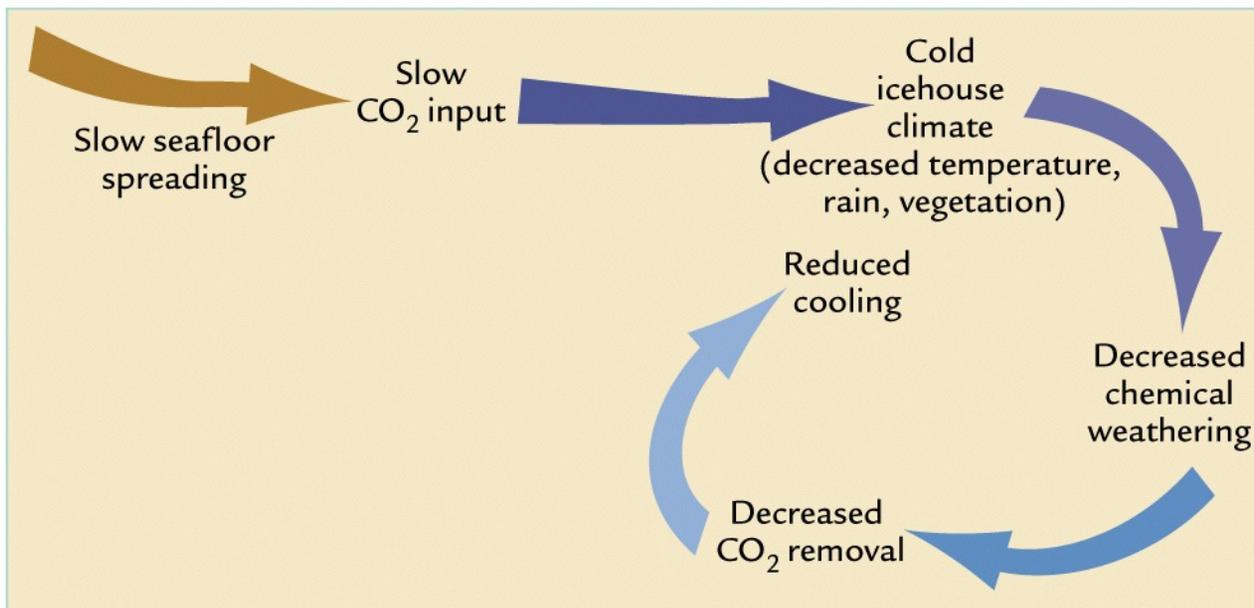
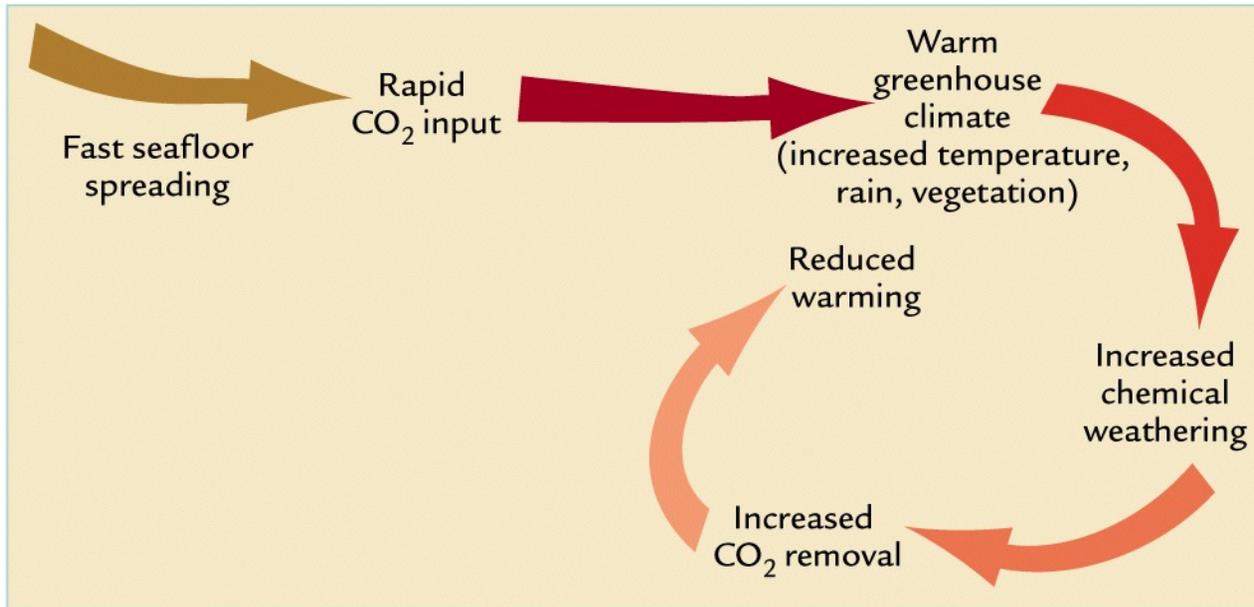


A

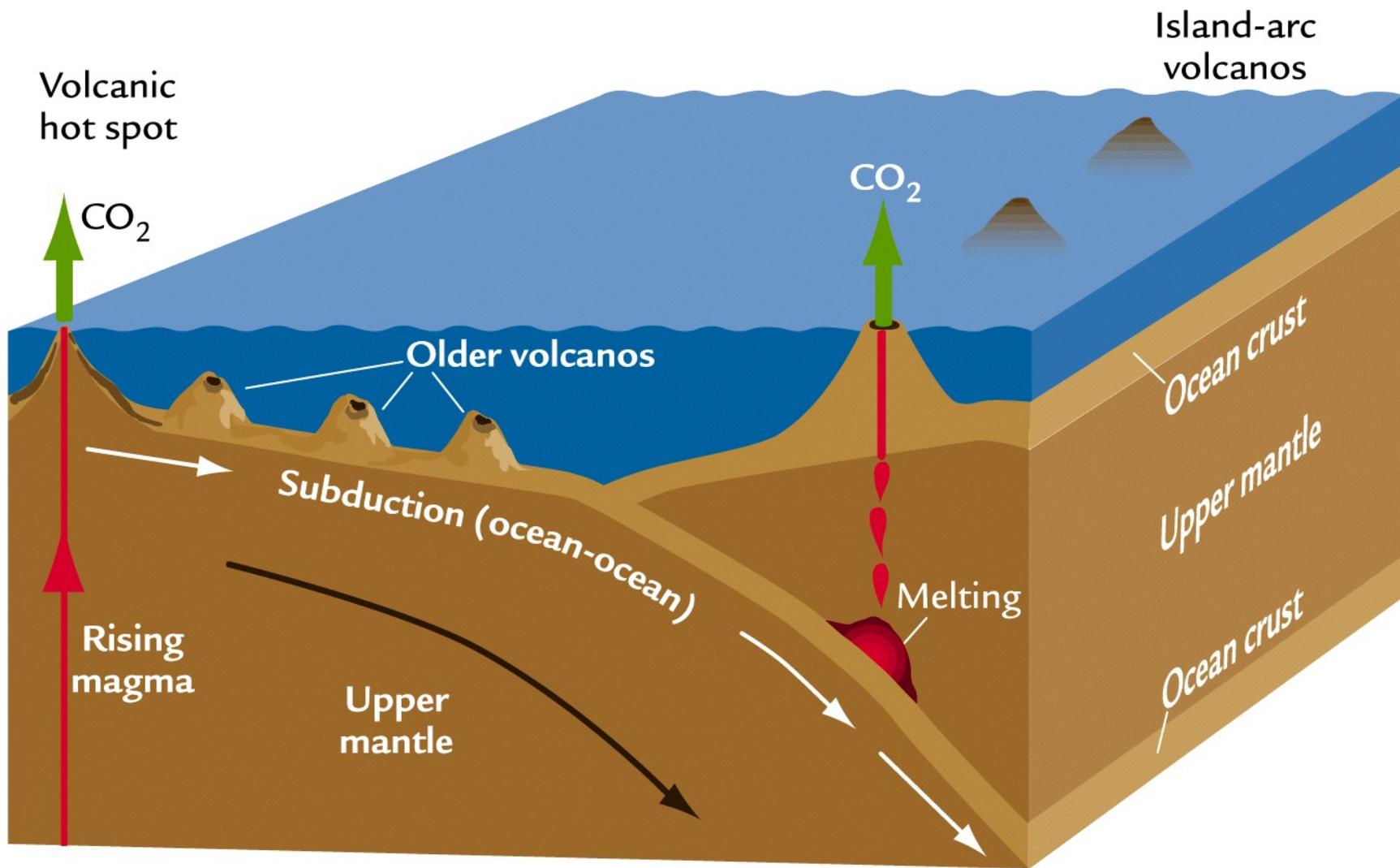


B

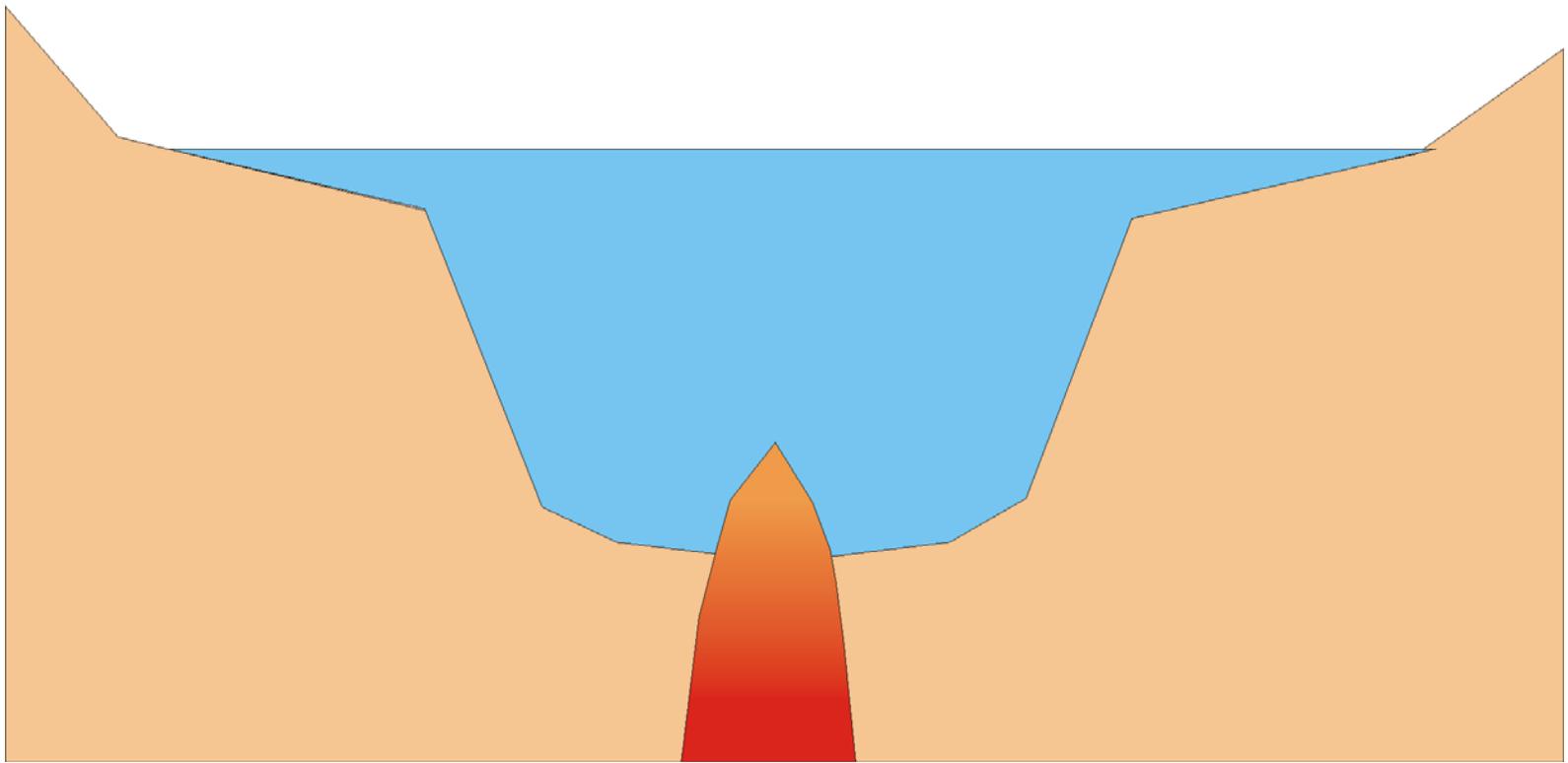


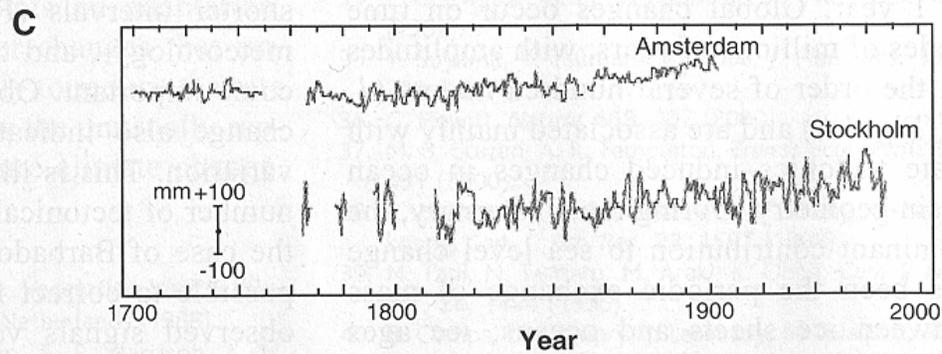
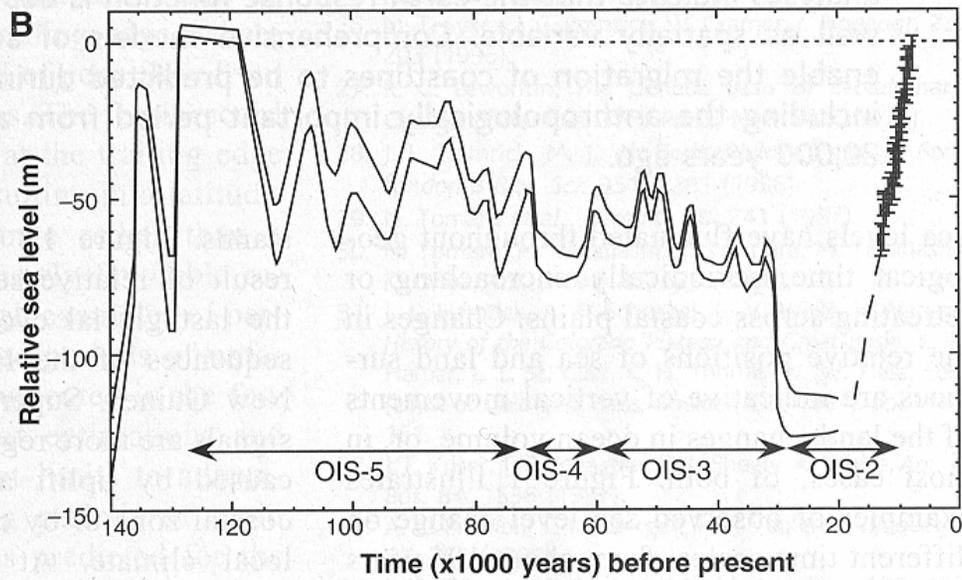
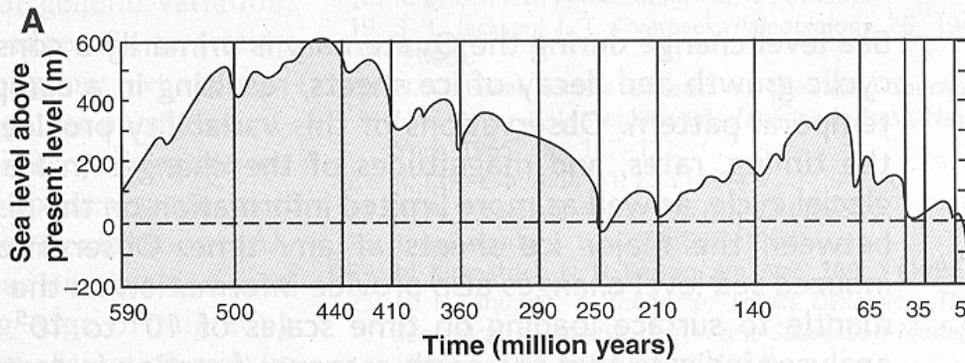


**Continental drift may cause climatic changes by changing the chemistry of the atmosphere**



Periods of relative high continental drift velocity



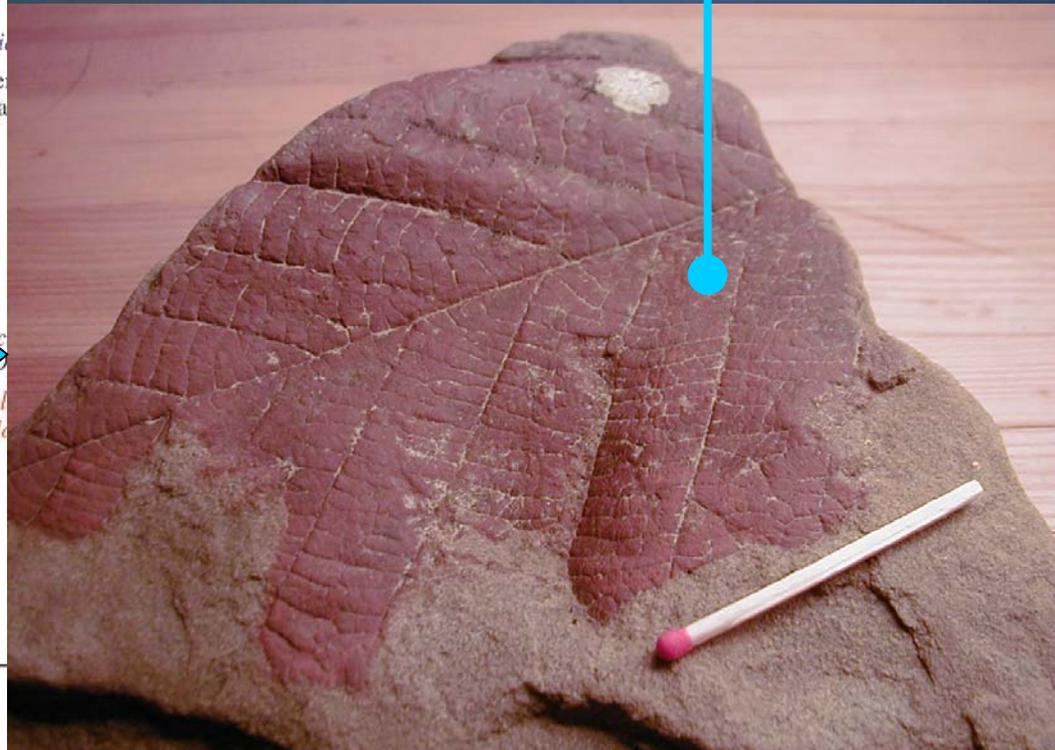
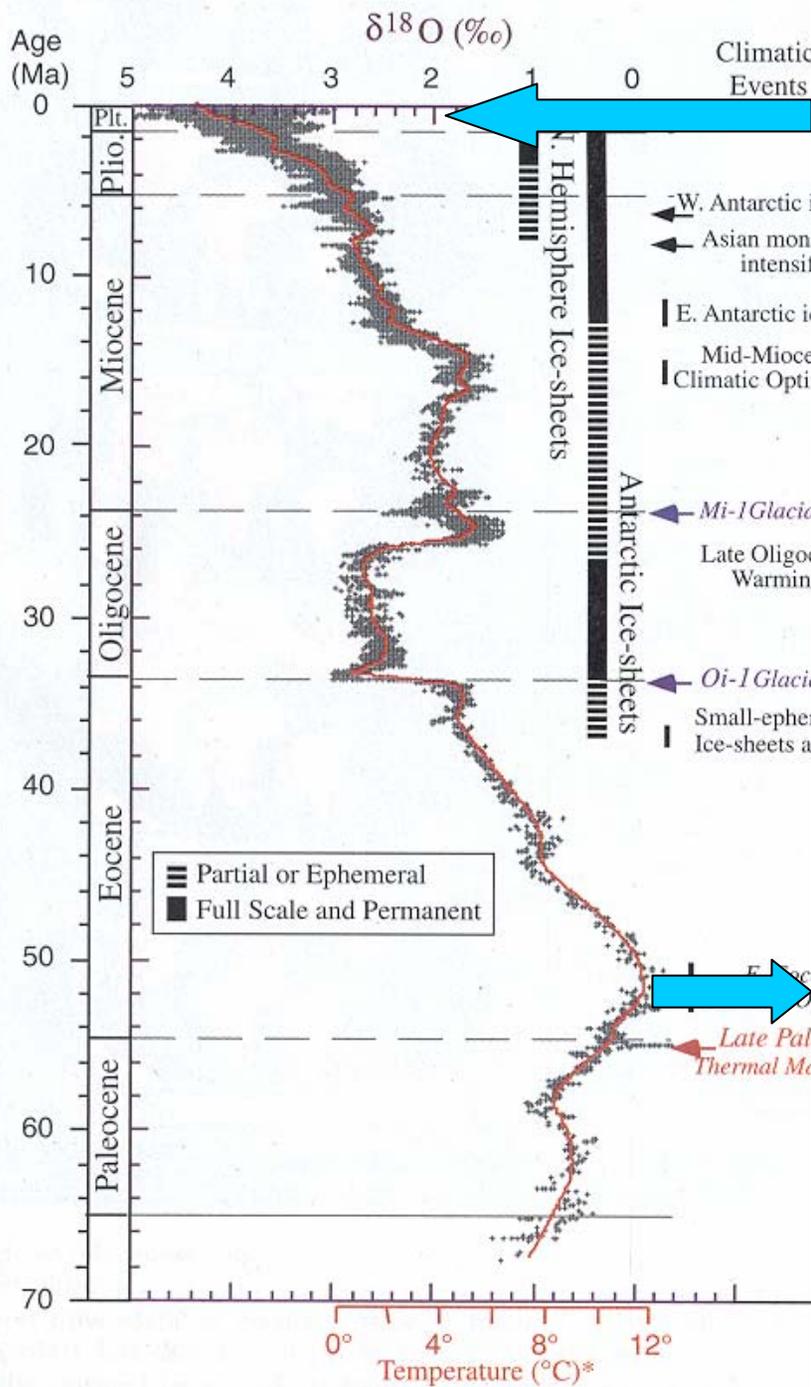


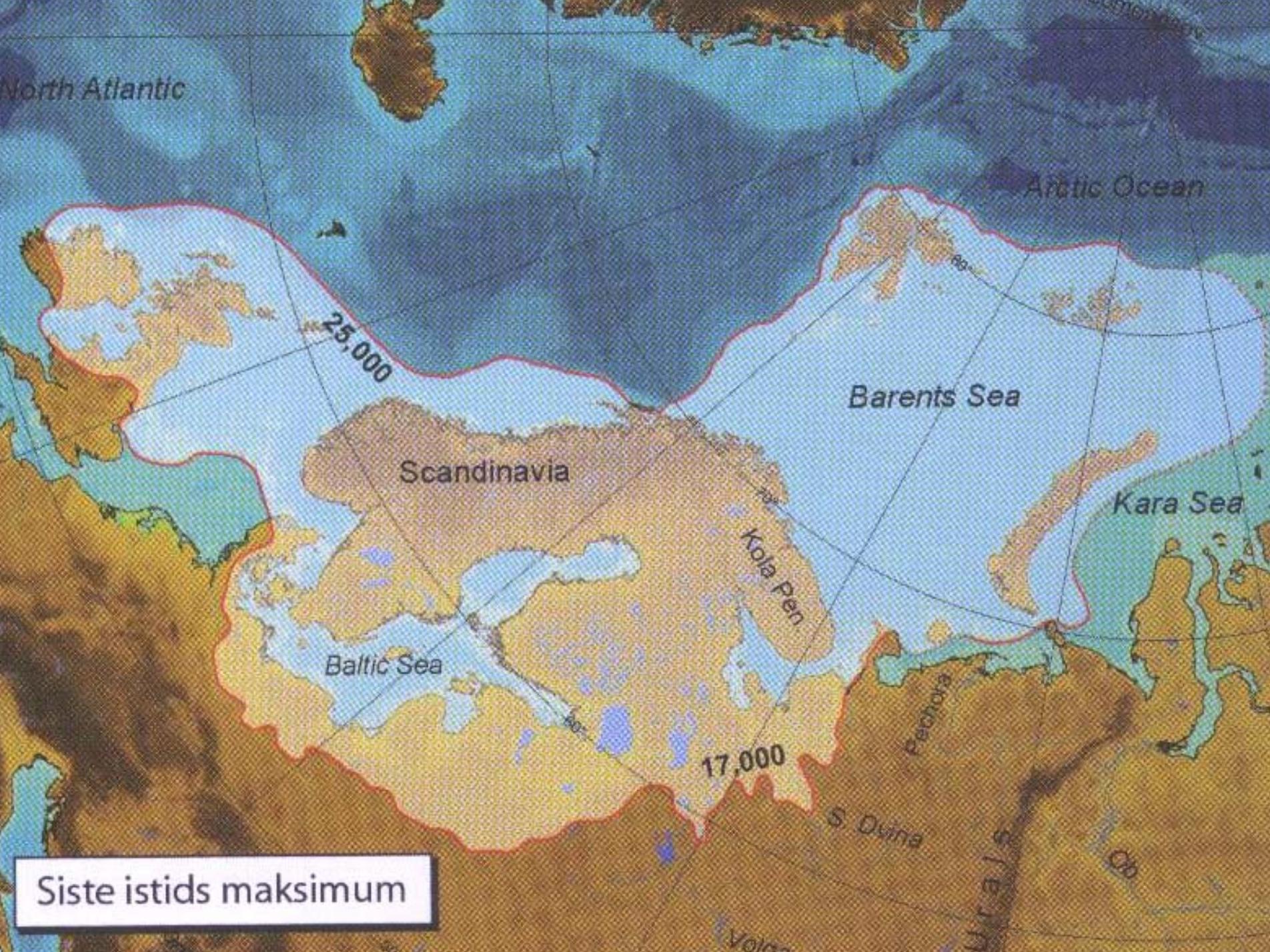
**1: Continental drift may cause sea level changes by changing the ocean bottom topography**

**2: Sea level changes may cause climatic changes by changing the land-sea configuration**



Tectonic history  
Continental drift  
Orbital changes  
Warm tertiary climate  
Quaternary glaciations  
Quaternary interglacials





North Atlantic

Arctic Ocean

25,000

Barents Sea

Kara Sea

Scandinavia

Baltic Sea

Kola Pen

17,000

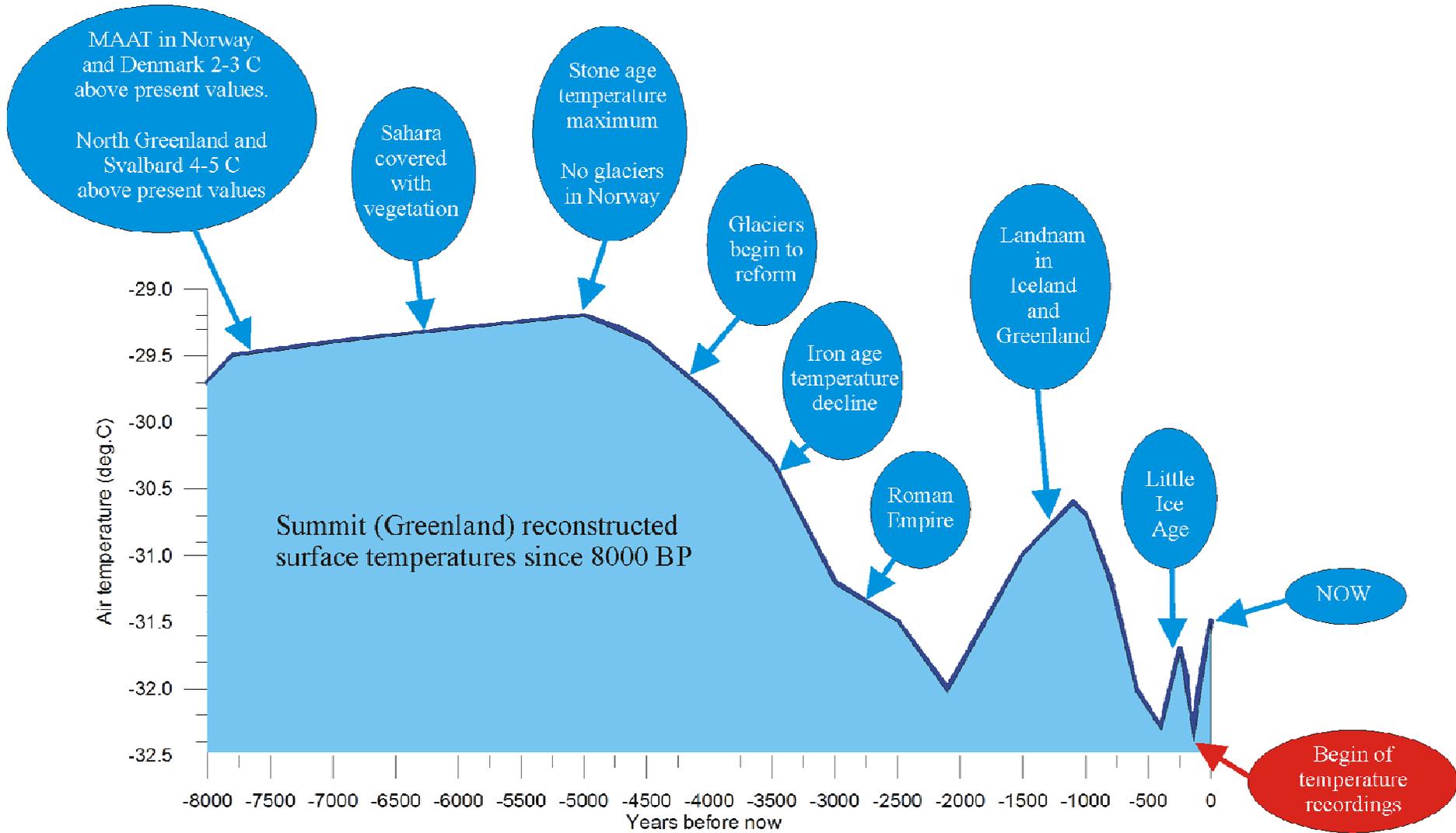
Pechora

S. Dvina

Ob

Ural

Siste istids maksimum







**Still on the road to learning !**